
GEOS S300 ULF DATA
and
GEOS AUXILIARY DATA

Decommutation, Calibration and Visualisation programs

Patrick Robert

Centre National de la Recherche Scientifique

Centre d'étude des Environnements Terrestre et Planétaires

Issue 1, revision 4

October 14, 1999

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Part I : GEOS S300 ULF DATA

1. Introduction

GEOS-1 and GEOS-2 spacecraft have provided, between April 1977 and January 1983, a lot of data, in particular in the ULF range (0.01-11.5 Hz). Initial decommutation program has been made by CNES in 1977, providing data in volts in the spinning system of the ULF antennas. Before any scientific exploitation, a deconvolution method, allowing the calibration of the ULF data in the spinning system of the satellite, and then in any geographic or magnetic system, has been developed in CRPE (Centre de Recherche en Physique des Environnements Terrestre et Planétaires). Then, using the information of the “Despin channels”, which gives information on the DC magnetic field in the spinning plane, a second method combining the Despin channels and the ULF channels has been developed in the years 1982-1984, allowing the comparison between the ULF search coil data and the high resolution DC magnetometer data. The document “deconvolution method used for calibration of the GEOS ULF data” (to be written) will describe in details these two methods. In this document, we present the “CALIULF” program which uses and applies these methods and which produces a “result” file containing in ascii code the ULF data at different levels of processing, the lowest level being the telemetry records, and the higher level being the combination of the ULF and Despin channels, calibrated in nT, and given in a fixed, non-spinning, geographic system.

This program, taking into account the two despin channels, is issued from a first version, written in 1982 in FORTRAN IV on the CDC NOS-BE 60 bits system, and then converted these last years in FORTRAN 77 running on Sparc stations with 32 bits words. The details on pre-existing different versions is given in comments in the FORTRAN code. This current version is now the only one existing program allowing the calibration of the ULF data, and remains the only one program to be maintain in the next years, in the frame of the “Centre de Données de la Physique des Plasmas” (CDPP), which now archives data of french experiments in Plasma Physics.

The “CALIULF” program needs 2 libraries for its use: First, the “GEOSDECOLIB” library, written now in FORTRAN 77 language, 32 bits system, which contains all the different modules necessary to extract the data contained in the binary raw data existing files and to put them into 32 bits standard words. Secondly, the “GEOSCALCLIB” library, which contains all the scientific modules necessary to compute and convert the data, at any levels of processing, for calibration and change of coordinate system. Details on these two libraries are given in annexe 1.

2. Binary raw data

2-a) Definition

The binary raw data consist now in binary files, made by CNES, which are an image of the original magnetic tape, with a standardized CCSDS date and time binary field preceding each data block. These data contain both the ULF data of the S300 search coil experiment, and the S331 magnetometer data, given with a low time resolution (88 seconds for GEOS-1, and 5.5 seconds for GEOS-2).

2-b) Accessibility

ULF raw data are available in the frame of the CDPP server. The list of all the archived files available from the CDPP is given in annexe 4.

Data can be requested on any time interval. The archived files will be then strictly cut to cover the specified time interval.

3. Decommuration and Calibration Program « CALIULF »

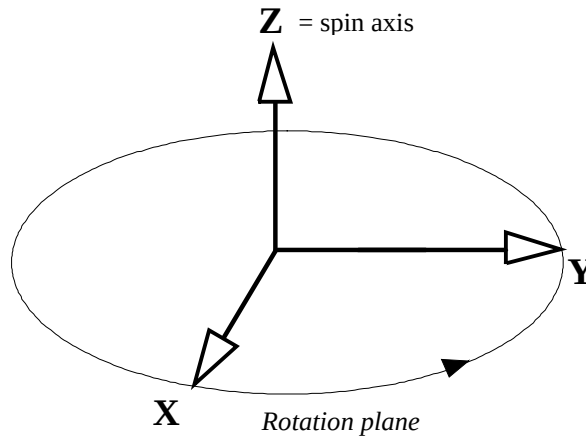
3-a) Role and Utility

The goal of the “CALIULF” program is to read the raw data binary files, extracts the physical data, converts them in volts, and apply the transfer function to obtain the ULF waves in nanoTesla (nT), by using a special deconvolution method (see the document “deconvolution method used for calibration of the GEOS ULF data”, to be written). These calibrated data are available in the spinning system of the satellites, or in a fixed system called “fixed meridian spacecraft”, where the Z axis is along the spin axis of the spacecraft, and the X axis is contained in the geographic meridian plane. Two other systems are available, the geographic VDH system, or the “magnetic strength line system” where the Z axis is along the magnetic field line defined by the DC magnetometer. All these systems are described in details in the next section.

A second level of processing combines the ULF data and the Despin channels data to obtain the upper level of the ULF data (see the document “deconvolution method used for calibration of the GEOS ULF data”). This method allows us to rebuild 2 components of the DC magnetic field, by using the amplitude and the phase of the two Despin x and y channels, roughly very close to a pure sine curve, and adding this information to the purely ULF data. These two components, which are included in the rotation plane, can be compared to the data of the S331 magnetometer experiments. The magnetometer data are available via the CALIULF program, but with a low time resolution (see section 3-c).

3-b) Description of the used coordinate systems

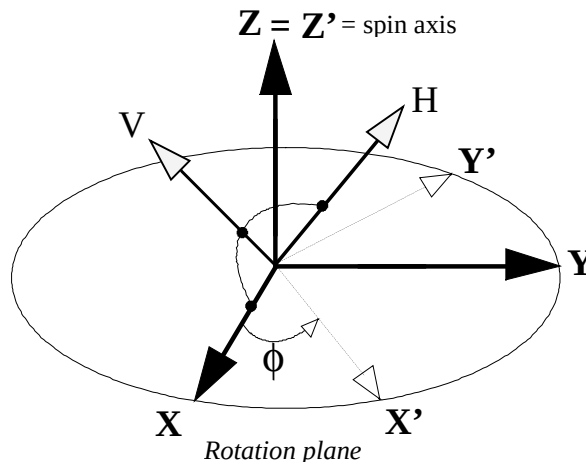
- *The “spinning” coordinate system:*



The **X,Y,Z** axis are the axis of the 3 ULF antennas, which are very close from the geometric axis of the spacecraft (less than 0.1-0.5 degrees, see the precise values of the spin axis tilt in the auxiliary data).

The **X** and **Y** axis define the rotation plane of the spacecraft.

- *The “fixed meridian spacecraft” coordinate system:*



The **Z**-axis is the spin axis of the spacecraft;

The **X'** and **Y'** axis are the rotating antennas, and define the rotation plane; there are the axis of the spinning coordinate system defined above.

The **V**-axis is the outwards local vertical, from the Earth to the spacecraft.

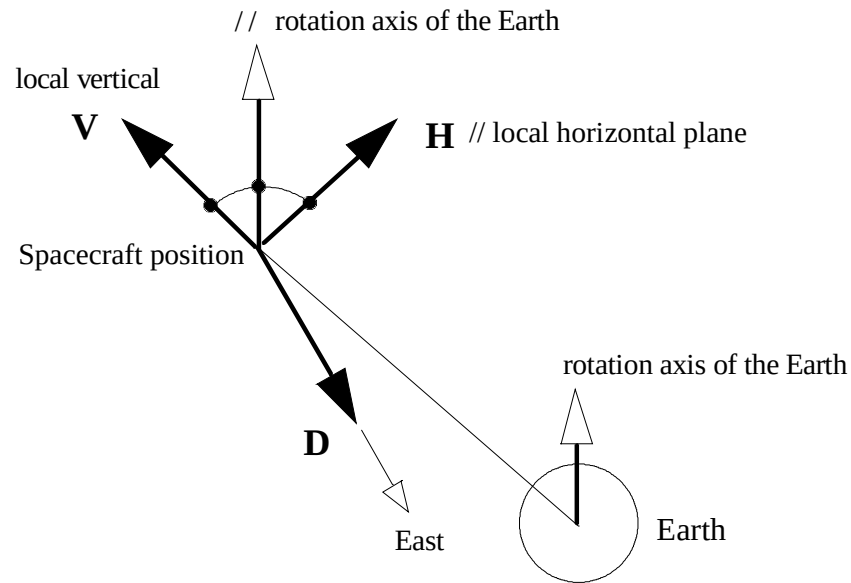
The **H**-axis is parallel to the horizontal local plane, positive to the North.

The **V-H** plane is a geographic meridian plane. See the more complete definition of the VDH system below.

ϕ is the azimuthal angle of the **X'** axis, and varies as $2\pi f_s t$

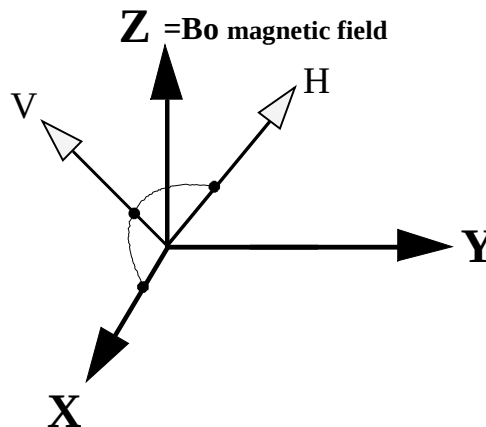
The **X**-axis correspond to the position of the **X'**-axis at the time where **X'** cross the geographic meridian plane, outward.

• **The VDH coordinate system:**



The V-axis is the outwards local vertical, to the spacecraft.
 The H-axis is parallel to the horizontal local plane, positive to the North.
 The V-H plane is a geographic meridian plane. The D-axis is azimuthal, eastwards.

• **The Magnetic Strength Line coordinate system:**



The Z-axis is along the DC local magnetic field;
 The V-axis is the outwards local vertical, from the Earth to the spacecraft.
 The H-axis is parallel to the horizontal local plane, positive to the North.
 The V-H plane is a geographic meridian plane.

The X-axis is perpendicular to the Z axis, and included in the local geographic meridian plane, outward.

All these system are local coordinate systems, which are dependent of the position of the spacecraft from the Earth.

3-c) Description of the caliulf.in file containing the input parameters

The CALIULF program reads a file, named caliulf.in, localised in the working directory, and containing the input necessary parameters for its execution. An example of this file is given below.

Example of caliulf.in file:

```
/data1/ida/robert/geos/geosdata/UBF_19780827_011043
28 08 1978      # date day month year
 9 14 40        # hour minute second
44              # duration in second (may be rounded)
1024            # number of points for the wave forms (2**N, 512 - 65 536)
0.1 11.5        # frequency limits, Hz, for the filtration
0              # output status: 0 normal, 1 for details on caliulf.out file
*              # panel status: * for all the panels
                (then visualisation by visuulf.pro program available)
                # or list of selected panels such as 6,10,14,15
```

The first line indicates the localisation of the binary ULF data file. The two following lines correspond to the asked date and time corresponding to the beginning of the analysed period. The beginning time must be rounded, because the ULF data are packed by a group of 4 Low Speed Format (LSF) corresponding to 4 x 1.37601 seconds, so the start time will be rounded around 5.504 seconds.

The duration parameter corresponds to the total period to be analysed, while the number of points of the wave forms corresponds to the window of the signal where the Fast Fourier Transforms (FFT) will be applied to process the signal and correct the effects of the non linear transfer function of the antennas. It is also the period of the data which will be written on the caliulf.resu file (see next sections). One must notice that over 1/16 of this window, on each side, a weighting function is applied, so the two short portions of time at the beginning and at the end of each wave form could be not fully significant (see the document “deconvolution method used for calibration of the GEOS ULF data”). The time sampling of the wave forms being 43.0003 millisecond, a wave form of 1024 points corresponds to 44.032 seconds, the minimum being 5.504 seconds (N=128) and the maximum being 2818.07, as around 47 minutes (N= 65536). Note that the elimination of the sine signal at the spin frequency is impossible for N< 512. In case of a given N less than 128 or greater than 65536, the 128 or 65536 values are assumed.

If the total duration asked is greater than the duration of the wave forms, defined by N, a next window or sequence is taken into account, and the true total duration corresponds to the number of taken wave forms, such as $N_w = (\text{asked duration}) / (\text{wave form duration})$.

The frequency limit parameters are used by the deconvolution method, which involves FFT, frequency filtration and change of coordinate system, taking into account that the value of the spin frequency is inside the ULF frequency range. Frequency bounds are given in the spinning coordinate system, and F1 cannot be less than 0.1 Hz, because the transfer function of the ULF antennas is equal to zero at F=0. The given values in this example cover the maximum that we can expect of these data. So the user can choose values inside the [0.1 – 11.5] interval, but not outside.

Output status: 0 is recommended, to have a limited size of the caliulf.out file. If the execution of the program shows some anomalies, the value of 1 will give much more information on the .out file, in particular for all the sequences.

There exists 14 levels of processing. The last parameter corresponds to the different required levels of processing. These levels are explained in the caliulf.resu section. The symbol “ * ” will leave to have in the caliulf.resu file all the 14 levels of ULF processing, and the 2 levels corresponding to the magnetometer data. This option is necessary for using the visuulf.pro IDL procedure which visualise all these data in a standard graphical form (see section 4). An example of these plots is given in annexe 4, where one can see the definition and the difference of the different levels of processing or different coordinate systems used.

If the user does not need the standard visuulf.pro procedure, does not want the intermediate levels of processing, and wants only for example the highest level of ULF data, and the corresponding magnetometer data, he only set this parameter to “ 14,15 ”, and will get into the caliulf.resu file only these two kinds of data. Then, he could use his own procedure to read the .resu file, visualise it, or apply his own process (see the description of the caliulf.resu file in next section).

3-d) Example of caliulf.out file describing the execution of the run

An example of caliulf.out file is given in annexe 2. This file gives precise informations about the data, the exact time of the TM formats, the position of the antennas and the DC magnetic field in different coordinate systems, the sense of rotation of the spacecraft, etc. It also shows the different steps of the data processing and is useful for checking the correct execution of the program. Since the CALIULF program and its libraries have been developed long years ago mainly by the author or other Frenchmen, in the CRPE and then the CETP, French laboratories, all the output comments are in French, and we are very sorry to present in the file the results in this language.

3-e) Description of the caliulf.resu file containing the calibrated data

An example of the caliulf.resu file is given in annexe 3. This file contains the results of the caliulf.exe program. It is in ascii format, to be commonly used by any software, in any language, on any machine. This file contains all the ULF data, corresponding to a given date and time period, with 14 levels of processing. It is completed by the DC magnetometer data, in two coordinate systems.

• Header:

Titles, input parameters deduced from the caliulf.in file, number of GEOS (1 or 2), number of sequences, selected panels (1 for yes), starting date and time. It is completed by the direction of the antennas in the VDH system at the starting time, the spin frequency and period, and particularly the coordinate in the VDH of the geographic meridian system used in many panels.

• data sets:

The header is followed by 16 groups of data, corresponding to the 16 panels plotted by the visuulf.pro procedure (see section 4 and annexe 4). Each set of data contains a sub-header with the time resolution (sampling period), the number of curves of the group or panel, and

the number of points. This sub-header is necessary for writing any reading program, and is used by the visuulf.pro graphical procedure. The description of these 16 groups is given below (see also annexe 4).

1) *XYZ Telemetry data:*

First level of the ULF data, in TM units, in the [0-255] interval. Time resolution is always 0.0430003 seconds

2) *Gains of the pre-amplifiers of the antennas:*

Two curves. The X-Y gains, in the spinning plane, are the same. The Z gain can be different. The small rounded symbol indicate the time resolution (always 1 Low Speed Format, so 1.37601 second).

3) *XYZ data gain-corrected:*

Telemetry data, after gain correction. Data in volts, in the spinning coordinate system of the antennas. $\Delta t = 0.0430003$.

4) *XYZ data gain-corrected, in SC fixed system:*

Same, but in a fixed coordinate system, where Z is along the Z antenna, and X is in the geographic meridian (see section 3-b). $\Delta t = 0.0430003$.

5) *XY despin channels in spinning system:*

The X and Y despin channels correspond to the main part of the DC magnetic field signal which is induced by the rotation of the X-Y antennas in the local DC magnetic field and detected by the search coils. This signal is removed by the hardware system from the initial signal of the antenna, in order to reduce the dynamic of the telemetry. The despin signals are mainly formed by a sine signal at the spinning frequency. Data in volts, with $\Delta t = 0.0430003$.

6) *XY despin channels in SC fixed system:*

Same, but in a fixed coordinate system, so the signals correspond roughly to the 2 DC magnetic field components of the DC magnetic field in this system.

7) *XY ULF + despin data in spinning system:*

The ULF and despin data are added, to reconstruct the full initial signals of the antennas. Data in volts, with $\Delta t = 0.0430003$.

8) *XY ULF + despin data in SC fixed system:*

Same, but in a fixed coordinate system, so the signals correspond to the 2 DC magnetic field components and the ULF data in this system. Data in volts, with $\Delta t = 0.0430003$.

9) *XYZ calibrated data, in nT, in spinning system:*

These data correspond to the #3 data level after deconvolution for removing the effect of the transfer function of the antennas. The data correspond to the ULF range only, in nT.

10) *XYZ calibrated data, as 9, in nT, in SC fixed system:*

Same, but in a fixed coordinate system. Since the Z component is already fixed, only the X-Y components are given.

11) XYZ calibrated data, in nT, in VDH system:

Same as 9, but in the VDH coordinate system. Each component is a mixture of the 3 others in the preceding system.

12) XYZ calibrated data, in NT, in strength line system:

Same as 11, but in the magnetic strength line coordinate system (see section 3-b). The DC magnetic field is deduced from the magnetometer data, and is kept constant during the whole duration of the wave form (over all the N points).

13) XYZ ULF+despin data, in nT, in spinning system:

These data correspond to the #7 data level where the deconvolution method is applied, to correct the effect of the transfer function of the antennas, both on the ULF range and the despin channels. A sine signal, at the spin frequency, with amplitude and phase computed from the whole signal, is subtracted to avoid a saturation of the signals in the spinning system by the sine induced by the strong DC magnetic field.

14) XYZ ULF+despin data, in nT, in SC fixed system:

Same as 13, but in the fixed system. The sine signal at spin frequency is taken into account before the change of coordinate, so the signal is not centered around zero, and have a significant offset, corresponding to the DC magnetic field. These data correspond to the highest level of processing, since we have in a fixed system the calibrated ULF data, with the two X and Y components of the DC magnetic field. These data can be compared to the DC magnetometer data, given in the next panels.

15) XYZ DC magnetic field, in nT, in SC fixed system:

These data are deduced from the magnetometer data, and computed in the same SC fixed coordinate system. The time resolution is 4 LSF, so 5.50404 seconds. A comparison with some sample of the high time resolution data of the magnetometer has shown a good agreement between the X-Y components of the magnetometer and the level 14 of the ULF data, which have nevertheless a more higher frequency range.

16) XYZ DC magnetic field, in nT, in VDH fixed system:

Same as 15, but in the VDH system.

3-f) Instruction for use

• **Installation on a new machine**

For the installation of the software on a new machine, for instance in the CDPF directory, the following directories and files are requested:

CDPF/geoslibs:

geosdecolib.f
geosdecolib.make
geoscalclib.f
geoscalclib.make

The execution of the shell scripts geosdecolib.make and geoscalclib.make will produce the object files geosdecolib.o and geoscalclib.o. The *.make files are given below ; a Fortran 77 compiler is requested.

geosdecolib.make:

```
echo
echo compil de geosdecolib.f, creation du .o
echo
f77 -c -C -ansi geosdecolib.f
```

geoscalclib.make:

```
echo
echo compil de la geoscalclib.f, creation du .o
echo
f77 -c -C geoscalclib.f
```

CDPP/caliulf:

caliulf.f
caliulf.make

The execution of the shell script caliulf.make will produce the object file caliulf.o and the executable file caliulf.exe. The caliulf.make file is given below, the F77 compiler is requested.

caliulf.make:

```
echo compil de caliulf.f
f77 -c -C caliulf.f
echo
echo link de caliulf.o avec les bibli geos
f77 -C caliulf.o ../geoslibs/geosdecolib.o \
    ../geoslibs/geoscalclib.o -o caliulf.exe
```

• *Running the caliulf.exe program*

After installation as above, the caliulf.exe program must be existing. The caliulf.in file must be included for providing to the program all the input parameters (see section above). It is useful, rather than executing directly the caliulf.exe program, to run the shell script caliulf.bat program, whose content is listed below:

caliulf.bat:

```
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo caliulf.bat : Calibration of GEOS ULF data
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
echo "Running caliulf.exe with parameters of caliulf.in"
echo "-----"
echo
echo "Please wait ..."
```

caliulf.exe

```
echo
echo "Proceedings of execution are in caliulf.out"
echo "Results are in caliulf.resu"
echo "Visualization available with visuulf.bat"
echo
echo "end of caliulf.bat"
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
```

This shell script file allows the user to get some comments during the execution of caliulf.exe. After running, the following files will occur :

caliulf.out	→ describes the execution of caliulf.exe (see section 3-d above)
caliulf.resu	→ contents the results of the different data processing asked (see section 3-e above)

4. Visualisation program « VISUULF »

4-a) Role and Utility

When the user indicates a “ * “ in the caliulf.in file, he ask for having all the data in the caliulf.resu file. These data correspond to the 14 levels of ULF data processing and the 2 levels of DC magnetic field indicated in the description of the caliulf.resu file (see section above). In order to have a complete view of the different levels and coordinate systems, it is useful to dispose of a standard graphical procedure, which produce directly a plot of all these data. It is the role of the visuulf.pro IDL (or PV-Wave) procedure, which reads the caliulf.resu file and produces groups of 4 PostScript files, named visuulf_s1_f1.ps, ... visuulf_s1_f4.ps, corresponding to the 4 frames of the sequence number 1. If N sequences are present in the caliulf.resu file (see above the description of the caliulf.in file), N groups of visuulf_SN_fM.ps are produced (M=1,4). This procedure works only with the “ * “ option in the caliulf.in file.

4-b) Example of produced visualisation

Annexe 4 gives an example of the print-out of the 4 PostScript files created by the visuulf.pro procedure. Each panel corresponds to the different levels generated in the caliulf.in file. The title of each panel indicates the type of plotted data, and the coordinate system used.

4-c) Instruction for use

- **Installation on a new machine**

The following files must be simply existing:

CDPP/caliulf/visuulf.pro
CDPP/caliulf/visuulf.bat

- **Running the visuulf.pro procedure**

The machine must have a IDL or a PV-Wave license. The visuulf.pro PV-Wave procedure is running by executing the visuulf.bat shell script given below:

visuulf.bat:

```
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo visuulf.bat : Visualization of GEOS ULF data
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
echo "File caliulf.resu must be existing"
echo "(this file is created by caliulf.bat)"
echo
#
echo "Running PV-Waves for data visualization"
echo "-----"
echo
echo      wave      <<!!
.run visuulf.pro
visuulf
caliulf.resu
!!
echo
echo "end of visuulf.bat"
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
```

The visuulf.pro procedure reads the caliulf.resu file, and plots the corresponding data into PostScript files.

Part II : GEOS AUXILIARY DATA

1. Introduction

GEOS-1 and GEOS-2 spacecraft have provided, between April 1977 and January 1983, a lot of data, for many different experiments, such as the ULF data described in part I. For all these experiments, the named “auxiliary data”, or DAVL meaning in french “Données Auxiliaires Variant Lentement”, consists in a set of data available for all the experiments. CNES has recently created a single file for all the DAVL of each satellite. In this part, we present the “CALIAUX” program which read these data, converts them into physical values, and produces a “result” file containing in ascii code these data.

The actual version of this program is unique for providing the auxiliary data, converted in physical units, and given in ASCII form. It remains the only program to be maintained in the next years, in the frame of the “Centre de Données pour la Physique des Plasmas” (CDPP), directed by CNES and CNRS, which now archives data of the Space Plasma Physics.

2. Binary raw data

2-a) Definition

The binary raw auxiliary data consist now in binary files, which are an image of the original magnetic tape, in which the original date and time fields have been replaced by CCSDS standardized date and time binary fields. A description of these binary auxiliary data is available in the documentation section of the CDPP server.

Description of the ASCII format of these data, as an output of the CALIAUX software, is given in section 3-e.

2-b) Accessibility

Auxiliary raw data are available in the frame of the CDPP server

The user is allowed to request these data for any time interval. He can also request for both the experiment data and the auxiliary data over the same time interval.

3. Decommuration and Calibration Program « CALIAUX »

3-a) Role and Utility

The goal of the “CALIAUX” program is to read the raw data binary files, and taking into account the original format, to transform them into classical physical units, and writing them into an ASCII file.

3-b) Description of the caliaux.in file containing the input parameters

The CALIAUX program reads a file, named caliaux.in, localised in the working directory, and containing the necessary input parameters for its execution. An example of this file is given below.

Example of caliaux.in file:

```
/data1/ida/robert/geos/geosdata/DAVL-GEOS1
21 12 1977      # date dd mm yy
00 01 30       # time hh mm ss
03.0           # duration in days
```

The first line indicates the localisation of the binary Auxiliary data file. The two following lines correspond to the asked date and time corresponding to the beginning of the analysed period. The beginning time must be rounded, because the Auxiliary data are packed by groups of 64 Low Speed Format (LSF) corresponding to 64 x 1.37601 seconds, so the start time will be rounded around 88.065 seconds. In order to avoid a starting period just before the asked day, it is recommended to set the start time to 00:01:30 second.

The duration parameter corresponds to the time interval to be calibrated, in decimal days. If the user wants to use the standard graphical visuaux.pro procedure (see section 4), which plots these data day by day into a PostScript file, this time interval must be an integer. If the user does not needs it, and wants to get only the caliaux.resu file on a given period, he can use a decimal day. Then he could use his own procedure to read the caliaux.resu file, visualise it, or apply any process (see the description of the caliaux.resu file in next section).

3-c) Example of caliaux.out file describing the execution of the run

An example of caliaux.out file is given in annexe 6. This file gives some information about the data, the exact time of the TM formats, the sense of rotation of the spacecraft, the number of blocks, etc. The first and last block of the given period is printed. Since the CALIAUX program and its libraries have been developed long years ago mainly by the author or other Frenchmen, in the CRPE and then the CETP, French laboratories, all the output comments are in French, and we are very sorry to present in the file the results in this language.

3-d) Description of the caliaux.resu file containing the calibrated data

An example of the caliaux.resu file is given in annexe 7. This file contains the results of the caliaux.exe program. It is in ascii format, to be commonly used by any software, in any language, on any machine. This file contains all the Auxiliary data, corresponding to a given date and time period, converted into physical units. The VDH coordinate system used for the vectors data is described in part I, § 3-b.

The header of the file is made by titles, by input parameters deduced from the caliaux.in file, by number of GEOS (1 or 2), number of blocks, starting date and time. It is completed by the title of each parameters.

The header is followed by all the data, given by group of two lines, as described below:

- | | |
|----------------------|---|
| 1) <i>nrec</i> : | Number of records read from the binary file.
Into the caliulf.resu file, the parameters corresponding to one record is written on 2 lines. |
| 2) <i>nJul</i> : | Number of Julian days from 1/1/1950 (1/1/1950 → <i>nJul</i> =0). |
| 3) <i>millisec</i> : | Millisecond of the day. |
| 4) <i>num lsf</i> : | Number of Low Speed Format (LSF).
Must be a multiple of 16. |
| 5) <i>lat</i> : | Latitude of the Spacecraft in degrees
(accuracy: 1/100 degree). |
| 6) <i>long</i> : | Longitude of the Spacecraft in degrees
(accuracy: 1/100 degree). |
| 7) <i>R km</i> : | Distance of the Spacecraft to the center of the Earth, in km
(accuracy: 10 km). |
| 8) <i>sunsat</i> : | Angle between the two lines Earth-centre to satellite
and Sun to Earth-centre (i.e eclipse at zero angle), in degrees
(accuracy: 1/100 degree). |
| 9) <i>rasc</i> : | Right ascension of the spin axis tilt in spacecraft coordinate,
in degrees (accuracy: 1/100 degree). |
| 10) <i>decli</i> : | Declination of the spin axis tilt in spacecraft coordinate,
in degrees (accuracy: 1/100 degree). |
| 11) <i>Qua</i> : | Quality of the data (0=good). |
| 12) <i>SpinX</i> : | Component X of the spin axis in VDH system
(accuracy: for digits after point) |

- 13) *SpinYc*: Component Y of the spin axis in VDH system
(accuracy: for digits after point)
- 14) *SpinZ*: Component Z of the spin axis in VDH system
(accuracy: for digits after point)
- 15) *Dip_X*: Component X of the magnetic dipole axis in VDH system
(accuracy: for digits after point)
- 16) *Dip_Y*: Component Y of the magnetic dipole axis in VDH system
(accuracy: for digits after point)
- 17) *Dip_Z*: Component Z of the magnetic dipole axis in VDH system
(accuracy: for digits after point)
- 18) *Sun_X*: Component X of the Sun axis in VDH system
(accuracy: for digits after point)
- 19) *Sun_Y*: Component Y of the Sun axis in VDH system
(accuracy: for digits after point)
- 20) *Sun_Z*: Component Z of the Sun axis in VDH system
(accuracy: for digits after point)

3-e) Instruction for use

- ***Installation on a new machine***

For the installation of the software on a new machine, the following directories and files are requested:

CDPP/geoslibs:

- geosdecolib.f
- geosdecolib.make
- geoscalclib.f
- geoscalclib.make

The geoscalclib and geosdecolib are described in detail in part I of this document.

The execution of the shell scripts geosdecolib.make and geoscalclib.make will produce the object files geosdecolib.o and geoscalclib.o. The *make files are given below ; a Fortran 77 compiler is requested.

geosdecolib.make:

```
echo
echo compil de geosdecolib.f, creation du .o
echo
f77 -c -C -ansi geosdecolib.f
```

geoscalclib.make:

```
echo
echo compil de la geoscalclib.f, creation du .o
echo
f77 -c -C          geoscalclib.f
```

CDPP/caliaux:

caliaux.f
caliaux.make

The execution of the shell script caliaux.make will produce the object file caliaux.o and the executable file caliaux.exe. The caliaux.make file is given below, the Fortran 77 compiler is requested.

caliaux.make:

```
echo compil de caliaux.f
f77 -c -C caliaux.f
echo
echo link de caliaux.o avec les bibli geos
f77 -C caliaux.o ../geoslibs/geosdecolib.o \
                ../geoslibs/geoscalclib.o -o caliaux.exe
```

• *Running the caliaux.exe program*

After installation as above, the caliaux.exe program must be existing. The caliaux.in file must be present to indicate to the program all the input parameters (see section above). It is useful, rather than executing directly the caliaux.exe program, to run the shell script caliaux.bat program, the content of which is listed below:

caliaux.bat:

```
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo caliaux.bat : Calibration of GEOS auxiliary data
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
echo "Running caliaux.exe with parameters of caliaux.in"
echo "-----"
echo
echo "Please wait ..."
```

caliaux.exe

```
echo
echo "Proceedings of execution are in caliaux.out"
echo "Results are in caliaux.resu"
echo "Visualization by 24H period available with visuaux.bat"
echo
echo "end of caliaux.bat"
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
```

This shell script file allows the user to have some comments during the execution of caliaux.exe. After running, the following files will occur :

caliaux.out → describes the execution of caliaux.exe (see section 3-d above)
caliaux.resu → contents the results of the different data processing asked
 (see section 3-e above)

4. Visualisation program « VISUAUX »

4-a) Role and Utility

It is useful to dispose of a standard graphical procedure, which produces directly a plot of all the Auxiliary data. It is the role of the visiaux.pro IDL (or PV-Wave) procedure, which read the caliaux.resu file and produces a PostScript file containing all the plots of the auxiliary data, with a page per day.

4-b) Example of produced visualisation

Annexe 8 gives an example of the print-out of a PostScript files produced by the visiaux.pro procedure. Each panel corresponds to the different types of data existing in the caliaux.in file.

4-c) Instruction for use

- **Installation on a new machine**

The following files must be simply existing:

CDPP/caliaux/visiaux.pro
CDPP/caliaux/visiaux.bat

- **Running the visiaux.pro procedure**

The machine must have a IDL or a PV-Wave license. The visiaux.pro PV-Wave procedure is running by executing the visiaux.bat shell script given below:

visiaux.bat:

```
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo visiaux.bat : Visualization of GEOS auxiliary data
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
echo "File caliaux.resu must be existing"
echo "(this file is created by caliaux.bat)"
echo
#
echo "Running PV-Waves for data visualization"
echo "-----"
```

```
echo
  wave    <<!!
.run visiaux.pro
visiaux
caliaux.resu
!!
echo
echo "end of visiaux.bat"
echo
echo XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
echo
```

The visiaux.pro procedure reads the caliaux.resu file, and plots the corresponding data into PostScript file.

Annexe 1: GEOSCALCLIB and GEOSDECOLIB libraries

The GEOSCALCLIB and the GEOSDECOLIB libraries are written in FORTRAN77, with lower characters, and consist in a serie of subroutines and functions. Since it has been developed long years ago mainly by myself or other Frenchmen of CRPE and then CETP, a French laboratory, all the comments are in French, and we are very sorry to present here the internal documentation code in this language. In order to list and comment the code, each module (subroutine or function) contains a pack of 3 lines of comments, just after the declaration of the modules, such as this example given below:

```
C      subroutine vdhxyz(x,y,z,a1,a2,a3,v,d,h,bzero)
C
C      -----
C **  categorie: astronomie et changement de base
C **  objet    : passage du repere satellite tournant au repere vdh
C **  auteur   : P. Robert CRPE 1977-1984
C      -----
C
```

This block of comments allows to define the category of the subroutine : here the subroutine is a part of the category “astronomy and change of coordinate systems”, see below. The second comment indicates the role of this module : here the module allows the user to convert a vector of the spinning coordinate system to the VDH coordinates. The third and last comment gives the author, his origin, and the date of creation of the module. These conventions are used for the two libraries, even if the decommutation library has been re-written recently.

This block, always existing for each module, is used by an external documentation program, developed for the management of this kind of codes, which produces a summary of the documentation, and which can be useful for any further developments. A header summarises the size of the code (number or lines), the number of different types (subroutines or function), and the different categories of modules. Then the list of modules is given, for each category, with their i/o arguments, a short summary of their role and authors issued from the key comments, and the line position of the module in the code. The results are given above.

A1-a) GEOSCALCLIB library

Header

DOCMOD: DOCUMENTATION OF FORTRAN MODULE SOURCE

adapted to GEOS/UBF fortran codes, PR 96

file source: geoscalclib.f

nb. of lines : 3195
nb. of function : 3
nb. of subroutine: 70
nb. of program : 0

nb. of modules : 73

nb. of categories: 4

- 1 astronomie et changement de base
- 2 mathematiques & traitement du signal
- 3 manipulation de tableaux
- 4 depouillement specifique GEOS/UBF

list of modules for each category:

category: astronomie et changement de base (11 modules)

type	module	arguments	objet	auteur	ligne
subr	calheur	(mill,ih,im,is,imil)	calcul ih,im,is,imil depuis les millisec. du jour	P. Robert CETP 1998	264
subr	caljj50	(iday,imonth,year,njj50)	calcul le nombre de jour julien depuis le 1/1/1950= 0	P. Robert CETP 1996	288
subr	carsph	(r,tet,phi,x,y,z)	conversion de coordonnees spheriques en cartesiennes	P. Robert CRPE 1977-1984	706
subr	cdfdot	(year,idoty,imonth,iday)	calcule la date depuis l'annee et le jour de l'annee	P. Robert CRPE 1993	725
subr	cdfjj50	(njj50,imonth,iday,year)	calcule la date depuis le nb de jour julien 1/1/1950=0	P. Robert CRPE 1993	780
subr	cleapy	(year,ileap)	compute_leap_year with ileap=1 for leap year, 0 if not	P. Robert CETP 1996	962
subr	codoty	(imonth,iday,year,idoty)	compute_day_of_the_year with idoty=1 for january 1	P. Robert CETP 1996	1007
subr	rot	(x,y,tet)	rotation sur x et y de centre 0 et d'angle teta	P. Robert CRPE 1977-1984	2562
subr	sphcar	(x,y,z,r,tet,phi)	conversion de coordonnees cartesiennes en spheriques	P. Robert CRPE 1977-1984	2751
subr	vdhxyz	(x,y,z,a1,a2,a3,v,d,h,bzero)	passage du repere satellite tournant au repere vdh	P. Robert CRPE 1977-1984	2968
subr	xyzvdh	(v,d,h,a1,a2,a3,x,y,z,bzero)	passage du repere vdh au repere satellite tournant	P. Robert CRPE 1977-1984	3103

category: mathematiques & traitement du signal (9 modules)

type	module	arguments	objet	auteur(s)	ligne
func	gauss	(x,ymoy,sig)	calcule pour x la fonction de gauss g(x,ymoy,sigma)	P. Robert CRPE 1977-1984	81
func	porte	(x,x1,x2)	calcule pour x la fonction porte(x,x1,x2)	P. Robert CRPE 1977-1984	117
subr	apm180	(deg)	recadrage entre +/- 180 d'un angle en degres	P. Robert CRPE 1977-1984	200
subr	apobor	(s,nbp)	apodisation aux bornes d'une forme d'onde complexe	P. Robert CRPE 1977-1984	220
subr	bornint	(i,i1,i2)	borne une valeur entiere i entre 2 autres i1 et i2	P. Robert CRPE 1977-1984	247
subr	calsin	(sig,n,fe,fs,as,phi)	calcul d'une sinusoide contenue dans un signal	P. Robert CRPE 1977-1984	588
subr	fftpat	(s,n,ind)	fft directe ou inverse (ind=1/2) sans table de sinus	Cooley/Tuquey/Capellini/Robert 1980	2152
subr	modpha	(s,rm,rp)	calcule le module et la phase (degre) d un nb. complexe	P. Robert CRPE 1977-1984	2345
subr	sinuso	(a,f,phi,fe,x,n)	charge x(n) par une sinusoide (phi en degres)	P. Robert CRPE 1977-1984	2725

category: manipulation de tableaux (10 modules)

type	module	arguments	objet	auteur	ligne
subr	add1t	(a,t1,t2,n)	additionne 1 tableau pondere $t2(i)=a*t1(i)+t2(i)$	P. Robert CRPE 1977-1984	134
subr	add2t	(a,t1,b,t2,t3,n)	additionne 2 tableaux ponderes $t3(i)=a*t1(i)+b*t2(i)$	P. Robert CRPE 1977-1984	156
subr	addxt	(x,t,n)	additionne un reel a un tableau	P. Robert CRPE 1977-1984	178
subr	sd1t12	(s,ta1,ta2,nbp)	met le cmplx s(n) dans ta1(n) et ta2(n)	P. Robert CRPE 1977-1984	2648
subr	strady	(s,tra,y,nbp)	met le cmplx s(n) et le real tra(n) dans y(3,n)	P. Robert CRPE 1977-1984	2778
subr	t1dt2	(t1,t2,n)	met t1(n) dans t2(n)	P. Robert CRPE 1977-1984	2802
subr	t1t2ds	(ta1,ta2,s,nbp)	met ta1(n) et ta2(n) dans le cmplx s(n)	P. Robert CRPE 1977-1984	2824
subr	t1t2t3dy	(ta1,ta2,ta3,y,nbp)	met ta1(n), t2(n) et ta3(n) dans y(3,n)	P. Robert CETP 1996	2847
subr	ydstra	(y,s,tra,nbp)	met y(3,n) dans le cmplx s(n) et le real tra(n)	P. Robert CRPE 1977-1984	3148
subr	ydt1t2t3	(y,ta1,ta2,ta3,nbp)	met y(3,n) dans ta1(n),ta2(n) et ta3(n)	P. Robert CETP 1996	3171

category: depouillement specifique GEOS/UBF (41 modules)

type	module	arguments	objet	auteur	ligne
func	gepy	(x)	fonction module du gain de l antenne electrique e y	P. Robert CRPE 1977-1984	100
subr	calbxyz	(phi,fe,fspin,rotdeg)	calcul du champ statique dans le repere fixe meridien	P. Robert CETP 1997	334
subr	calgeos	(njd50,igeos)	calcule le num. du satellite (1 ou 2)	P. Robert CRPE 1977-1984	396
subr	calphi	(lsf1,teta,phi,lsfphi,fe)	calcul du teta et phi rafraichi pour un lsf1 donne	P. Robert CRPE 1977-1984	429
subr	calrot	(rotdeg)	calcul de l'angle de rot. entre x et le merid. geog	P. Robert CRPE 1977-1984	521
subr	calsinu	(x,n,fs,fe,a,pha,texte)	calcule ampli. & phase de la sinusoide de spin (calsin)	P. Robert CETP 1996	653
subr	caltemp	(njul50,milduj,jj,mm,ian,ih,im,is,ims)	calcul le temps du format courant	P. Robert CETP 1998	690
subr	chbavdh	(s,tra,nbp,fe,fspin)	chang. base formes d'onde sat. tournant au repere vdh	P. Robert CRPE 1977-1984	823
subr	chbaxyb	(s,tra,nbp,fe,fspin,tet,phi)	chang. base formes d'onde sat. tournant au repere xyb	P. Robert CRPE 1977-1984	870
subr	consaux	(nubloc,nbblo,njj,mil,lsf,njip,milp,lsfp,nbbm)	verifie et complete la consecutivite des blocs des DAVL	P. Robert CETP 1999	1056
subr	consecf	(lsf1,lsf2,lsfp,lsf11)	test de consecutivite des formes d'ondes	P. Robert CRPE 1977-1984	1117
subr	constim	(nubloc,nbblo,njj,mil,lsf,njip,milp,lsfp,nbbm)	verifie la consecutivite du temps des blocs des DAVL	P. Robert CETP 1999	1153
subr	correch	(s,nbp,fe,j)	correction des retards d echantillonnage des voies xyz	P. Robert CRPE 1977-1984	1203
subr	corrgea	(s,n,df,ip)	correction des grains d antenne sur tableau s(n)	P. Robert CRPE 1977-1984	1243
subr	deconvo	(s,tra,y,nbp,f1,f2,fe)	deconvolution des formes d'onde x y z	P. Robert CRPE 1977-1984	1285
subr	despin	(s,tra,n,fe,fs,idespi,ax,ay,az,phix,phiy,phiz)	despinage des formes d'onde s et tra (algo. demod)	P. Robert CRPE 1977-1984	1365
subr	ebdsvdh		ecriture des donnees du champ mag. Stat. dans le VDH	P. Robert CRPE 1977-1984	1491
subr	ebdsxyz		ecriture des donnees du champ statique ds XYZ sat.mer.	P. Robert CETP 1997	1531
subr	ecdavl	(nu,nb,jc,mil,lsf,ria,rlo,rkm,c,d,s,sts,as,de,ig)	ecriture sur fich. resu d'un bloc de davl	P. Robert CETP 1998	1571
subr	eentaux	(lgeos,indata,nbbloc)	ecriture sur . resu de l'entete aux. (date, positionn.)	P. Robert CETP 1996	1635
subr	eentubf	(lgeos,indata,nbseq)	ecriture sur . resu de l'entete UBF (date, positionn.)	P. Robert CETP 1996	1680
subr	efinpro	(com)	ecriture du status de fin normale pour le fichier .out	P. Robert CETP 1996	1731
subr	enumseq	(lwrite,nbseq)	ecriture du numero de sequence pour le fichier .out	P. Robert CETP 1996	1760
subr	etab2	(ta1,ta2,nbp,res,ipan,com)	ecriture sur fich. resu de ta1(nbp) et ta2(nbp)	P. Robert CETP 1996	1800
subr	etab3	(ta1,ta2,ta3,nbp,res,ipan,com)	ecriture sur fich. resu de ta1(nbp), ta2(nbp), ta3(nbp)	P. Robert CETP 1996	1858
subr	etabs	(s,nbp,res,ipan,com)	ecriture sur fich. resu de xy en volt rep. fixe	P. Robert CETP 1996	1917
subr	etabst	(s,tra,nbp,res,ipan,com)	ecriture sur fich. resu de y(3,nbp)	P. Robert CETP 1996	1975
subr	etaby	(y,nbp,res,ipan,com)	ecriture sur fich. resu de y(3,nbp)	P. Robert CETP 1996	2034
subr	etemseq	(njul50,jj,mm,ian,ih,im,is,ims)	ecriture sur fich. resu du temps de la sequence	P. Robert CETP 1998	2092
subr	fftdu	(s,nbp)	calcul une fft directe d un complexe s(n)	P. Robert CRPE 1977-1984	2120
subr	fftinu	(s,nbp)	calcul une fft inverse	P. Robert CRPE 1977-1984	2153
subr	flitond	(s,n,df,f1,f2)	filtrage dune forme donde	P. Robert CRPE 1977-1984	2255
subr	longut	(mot,nc)	calcul de la longueur utile d'une chaine de caracteres	P. Robert CETP 1998	2293
subr	longsb	(mot,nc)	calcul de la longueur sans blanc d'une chaine de carac.	P. Robert CETP 1998	2320
subr	mophafs	(cgfs,rfs,phafs)	module et phase de la fonction de transfert a fspin	P. Robert CETP 1996	2373
subr	psatmer	(rotdeg)	calcule les axes du repere meridien dans le VDH	P. Robert CETP 1997	2409
subr	psatvdh		calcule la direction des antennes dans le vdh au lsf1	P. Robert CRPE 1977-1984	2490
subr	rotfix	(s,nbp,fe,fspin,rotdeg)	passage d'un repere en rotation a un repere fixe	P. Robert CRPE 1977-1984	2586
subr	sensrot	(njul50,s)	calcule le sens de rotation du satellite	P. Robert CRPE 1977-1984	2672
subr	titpart	(nompar)	titre de la partie pour le fichier resultat .out	P. Robert CETP 1996	2872
subr	titcali	(nompro)	titre/entete de caluif pour le fichier resultat .out	P. Robert CETP 1998	2896
subr	titauxi	(nompro)	titre/entete de caliaux pour le fichier resultat .out	P. Robert CETP 1998	2933
subr	verifcb	(phi)	verifie les calculs de changement de base	P. Robert CRPE 1977-1984	3012

A1-b) GEOSDECOLIB library

This library has been many times re-written, since it contains the modules allowing the extraction of the data from the binary data files, and thus, depends on the hardware of the computer. It contains only one no standard Fortran 77 function, the "and" function, which is nevertheless available on most machines, such as SUN 32 bits workstation where this last version has been developed.

Header

```
-----  
DOCMOD: DOCUMENTATION OF FORTRAN MODULE SOURCE
```

```
adapted to GEOS/UBF fortran codes, PR 96  
-----
```

```
file source: geosdecolib.f
```

```
nb. of lines      : 1044
```

```
nb. of function   :    1
```

```
nb. of subroutine :    8
```

```
nb. of program    :    0
```

```
nb. of modules    :    9
```

```
nb. of categories:    1
```

```
    1 decommutation GEOS/UBF  
-----
```

list of modules for each category:

only one category: decommutation GEOS/UBF (9 modules)

type	module	arguments	objet	auteur(s)	ligne
subr	aconde	(nbp,y,ux,uy,uz,dx,dy,lsf1,lsf2,fe,fspin)	acquisition et decom. des donnees UBF	PR,MR,OCJ, refait PR 1996, revu 98	62
subr	detmls	(lsf,jc,mj,fe,ib,ig,idx,idy,bvdh,bz,tet,phi,a,fs)	lecture et decom. d'un bloc de 4 LSF TMLS	PR,MR,OCJ, revu PR 1996, reecrit 98	266
subr	dedavl	(jcn,mil,lsf,rla,rlo,rkm,c,d,s,sts,asc,dec,iq)	lecture et decom. d'un bloc de DAVL	P. Robert, CESTP 1998	490
subr	dedate	(natfic,ient,j,mill,its)	decommutation de la datation des blocs	O. Coeur-Joly CESTP 1995, refait PR 98	663
subr	lebloc	(ifc,nos,nol,itab,ieof)	lecture d'un bloc de donnees TMLS ou DAVL	P. Robert, CESTP 1998	717
subr	posfic	(natfic,njul50)	positionnement sur le fichier tmls ou davl	O. Coeur-Joly CESTP 1995, refait PR 98	748
subr	tramil	(mill,ih,im,is,imil)	transforme les millisec. du jour en ih,im,is,imil	O. Coeur-Joly CESTP 1995, verifie PR 98	977
subr	isigne	(mot, nbit)	inversion de signe si mot > 2**(nbit-1)	O. Coeur-Joly CESTP 1995, verifie PR 98	1000
func	ipicmo	(mot, nbit)	extrait les nbit premiers bits de mot	O. Coeur-Joly CESTP 1995, revu PR 98	1020

Annexe 2: Example of caliulf.out file

```
GEOS-1 / GEOS-2

CALIULF V 4.0

Decommutation et calibration de tous les parametres des donnees UBF

P.Robert, CETP, Juin      1996
      revu  Avril      1997
      revu  Février    1998
Version pour le CDPP, Decembre 1998
```

```
lfulfin:

nom du fichier de donnees UBF ? (path complet)
/data1/ida/robert/geos/geosdata/UBF_19780827_011043

Date (jj,mm,aa) ? [27,04,1977]
  28  8  1978
Heure de debut (hh,mn,ss) ? [15,00,00]
  9  14  40
Duree de la periode (nb entier de sec.) ? [90]
  40
Nb de points d'une F.O. (-> res. spec.) ? [2048]
  1024
Bornes en frequence / deconvolution ? [0.5,11.5]
  1.00000E-01  11.5000
Detail de chaque sequence sur .out (1=oui) ? [0]
  0
Panels a selectionner pour le .resu [6,10,14 ou *]
*

resume des parametres d'entree:
28 08 1978 09 14 40    40  1024   0.1  11.5 0 *

soit  1 sequences de    44.0323 sec.
-----
calgeos:

GEOS-1/2:  GEOS-2
-----
sensrot:

sens de rotation= electron
-----
posfic: positionnement sur fichier de type tmls

      debut fichier : njul50,ih,im,is      = 10465  1 10 43

      fichier correspondant a GEOS-2
      nombre de records:      32479
      taille en octets :    31699504

      fin fichier      :      = 10467  3  4 51
      demande         :      = 10466  9 14 40
      fait            : njul50,ih,im,is,imil= 10466  9 14 40 862
-----
eentubf:

Donnees UBF GEOS - Resultats de Caliulf.exe
parametre d'entree: 28 08 1978 09 14 40    40  1024   0.1  11.5 0 *
GEOS:  2
Nombre de sequences :    1
Status des panels retenus:  1  2  3  4  5  6  7  8  9 10 11 12 13 14 15 16
                           1  1  1  1  1  1  1  1  1  1  1  1  1  1  1  1
Nombre de panels:  16
```

```
=====
enumseq:   sequence numero 1 //////////////////////////////////////
-----
aconde:

tranche de signal numero 1
acquisition des formes d'onde x y z sur 1024 points
valeurs corrigees des gains de pre-ampli
unites de sortie: volts
nombre de groupe de 4 LSF: 8
 1 datation: 18523572 10466 9 14 40 862 Gxz| 25- 25| 25- 25| 25- 25| 25- 25
 2 datation: 18523576 10466 9 14 46 366 Gxz| 25- 25| 25- 25| 25- 25| 25- 25
 3 datation: 18523580 10466 9 14 51 870 Gxz| 25- 25| 25- 25| 25- 25| 25- 25
 4 datation: 18523584 10466 9 14 57 375 Gxz| 25- 25| 25- 25| 25- 25| 25- 25
 5 datation: 18523588 10466 9 15 2 879 Gxz| 25- 25| 5- 5| 5- 5| 5- 5
 6 datation: 18523592 10466 9 15 8 383 Gxz| 5- 5| 5- 5| 5- 5| 5- 5
 7 datation: 18523596 10466 9 15 13 887 Gxz| 5- 5| 5- 5| 5- 5| 5- 5
 8 datation: 18523600 10466 9 15 19 391 Gxz| 5- 5| 5- 5| 5- 5| 5- 5

premier lsf: 18523572      dernier lsf: 18523603
Duree: 44.03 sec.
-----
consec: verification de la consecutivite des formats dans la sequence 1
-----
ebdsvdh:

donnees magnetometre S331 dans le VDH

      lsf      Bv      Bd      Bh      Br      Btet      Bphi
18523572 -10.000  57.700 -95.900  112.366  148.590  99.832
18523576 -3.600  47.600 -97.500  108.559  153.914  94.325
18523580 -2.600  42.200 -109.600  117.472  158.905  93.526
18523584 9.000  35.600 -120.200  125.684  163.013  75.812
18523588 -7.000  39.600 -133.300  139.234  163.212  100.024
18523592 -33.300  37.600 -124.400  134.157  158.014  131.529
18523596 24.400 119.600 -114.900  167.635  133.268  78.469
18523600 -16.200  81.000 -107.100  135.255  142.358  101.310
-----
psatvdh:

position des antennes dans le vdh au format 18523572

      V      D      H      r      teta      phi
axe x= -0.80313 -0.59580 -0.00226 1.000  90.130 -143.430
axe y= 0.59580 -0.80312 -0.00393 1.000  90.225 -53.430
axe z= 0.00052 -0.00451 0.99999 1.000  0.259 -83.370
-----
calphi:

fspin= 0.169815 Hz  Tspin= 5.88877 s
teta= 148.888
phi=-116.583 au debut du lsf 18523572 calcule depuis phi=-132.068
      au lsf 18523572 avec To= 0.253E+00
      bphi(1)=-132.068 dphi= 15.485
-----
verifcb:

verification des changements de base au format 18523572

B xyz lu directement (phi non defini)      B xyz= -38.901 -43.101 -96.203
                                           B mod= 112.366
                                           teta = 148.888
                                           phi = -132.068

vrai phi rafraichi calcule par calphi: phi= -116.583

B xyz lu avec phi corrige (valeurs exactes)  B xyz= -25.982 -51.923 -96.203
                                           B mod= 112.366
                                           teta = 148.888
                                           phi = -116.583
```

B vdh lu directement B vdh= -10.000 57.700 -95.900
 B mod= 112.366
 teta = 148.590
 phi = 99.832

B xyz recalculé par vdh et les angles d'euler B xyz= -26.129 -51.921 -96.164
 B mod= 112.366
 teta = 148.850
 phi = -116.714

B vdh recalculé par xyz et les angles d'euler B vdh= -10.120 57.614 -95.939
 B mod= 112.366
 teta = 148.629
 phi = 99.962

angles d'Euler au format 18523572 a1,a2,a3= -150.060 0.260 6.630

 calrot:

les phases sont relatives a x dans le plan du meridian géographique
 correction= 143.430 deg.

 psatmer:

position des antennes dans le repere meridian
 (interpolée depuis le lsf1= 18523572)

	V	D	H	r	teta	phi
axe x.mer=	1.00000	0.00000	-0.00052	1.000	90.030	0.000
axe y.mer=	0.00000	0.99999	0.00451	1.000	89.742	90.000
axe z.mer=	0.00052	-0.00451	0.99999	1.000	0.259	-83.370

 calbxyz:
 changement de base sur B331 repere sat. tournant -> repere satellite fixe
 x meridian rotdeg= 143.4 nb4lsf= 8 fs= 0.16981

ebdsxyz:
 donnees magnetometre S331 dans XYZ sat/meridian

lsf	Bv	Bd	Bh	Br	Btet	Bphi
18523572	-10.069	57.181	-96.203	112.366	148.888	99.987
18523576	-3.720	47.024	-97.775	108.559	154.245	94.523
18523580	-2.749	41.625	-109.816	117.472	159.200	93.779
18523584	9.084	35.017	-120.365	125.684	163.272	75.457
18523588	-7.191	38.921	-133.490	139.234	163.485	100.468
18523592	-33.431	37.057	-124.528	134.157	158.160	132.055
18523596	24.253	119.080	-115.469	167.635	133.536	78.488
18523600	-16.444	80.464	-107.466	135.255	142.612	101.550

 mophafs:
 amplitude= 0.002280 phase= 89.163651

partie 1: telemesure

 etaby

1) XYZ Telemetry data
 resolution (s): 0.0430003
 Nb. de courbes: 3
 Nb. de points : 1024

etab2:
 2) Gains XY and Z
 resolution (s): 1.3760099
 Nb. de courbes: 2
 Nb. de points : 32

etab3:

3) XYZ volt, SC spinning system
resolution (s): 0.0430003
Nb. de courbes: 3
Nb. de points : 1024

rotfix:

changement de base sur les formes d'onde, repere final=repere sat. fixe
x meridien rotdeg= 143.4 nbp= 1024 fs= 0.16981

etabs:

4) XY volt, SC fixed system with X in geo. meridian
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

partie 2: UBF et Despins bruts

etab2:

5) DX,DY volt, SC spinning system
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

rotfix:

changement de base sur les formes d'onde, repere final=repere sat. fixe
x meridien rotdeg= 143.4 nbp= 1024 fs= 0.16981

etabs:

6) DX,DY volt, SC fixed system with X in geo. meridian
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

etab2:

7) X+DX,Y+DY volt, SC spinning system
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

rotfix:

changement de base sur les formes d'onde, repere final=repere sat. fixe
x meridien rotdeg= 143.4 nbp= 1024 fs= 0.16981

etabs:

8) X+DX,Y+DY volt, SC fixed system with X in geo. meridian
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

partie 3: UBF deconvolues

despin:

despinage des formes d'ondes, avec fs= 0.16981 et fe= 23.25565
amplitudes: 0.019808 0.019526 0.002689
phases (d): -42.055084 49.369930 61.343052

deconvo:

deconvolution des 3 formes d'onde- t.f. directe
fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre 0.100 et 11.500 Hz
correction du retard d'echantillonnage de la voie 1
fft inverse effectuee
fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre 0.100 et 11.500 Hz
correction du retard d'echantillonnage de la voie 2
fft inverse effectuee
fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre 0.100 et 11.500 Hz
correction du retard d'echantillonnage de la voie 3
fft inverse effectuee
t.f. inverse- retour au domaine temps

etabst

9) XYZ nT, SC spinning system
resolution (s): 0.0430003
Nb. de courbes: 3
Nb. de points : 1024

rotfix:

changement de base sur les formes d'onde, repere final=repere sat. fixe
x meridien rotdeg= 143.4 nbp= 1024 fs= 0.16981

etabs:

10) XYZ nT, SC fixed system with X in geo. meridian
resolution (s): 0.0430003
Nb. de courbes: 2
Nb. de points : 1024

chbavdh:

changement de base sur les formes d'onde, repere final = repere vdh

etabst

11) XYZ nT, VDH system
resolution (s): 0.0430003
Nb. de courbes: 3
Nb. de points : 1024

chbaxyb:

changement de base sur les formes d'onde, repere final = repere magnetique xyb

etabst

12) XYZ nT, XYB strength line system
resolution (s): 0.0430003
Nb. de courbes: 3
Nb. de points : 1024

partie 4: UBF + Despins deconvolues

calsinu:

spin de X amplitude= 0.019808 phase= -42.055 deg.

calsinu:

spin de Y amplitude= 0.019526 phase= 49.370 deg.


```
-----
calsinu:
spin de Z          amplitude=  0.002689   phase=   61.343 deg.
-----
calsinu:
spin de DX         amplitude=  0.109538   phase=  -60.162 deg.
-----
calsinu:
spin de DY         amplitude=  0.108372   phase=   28.370 deg.
-----
calsinu:
spin de X+DX       amplitude=  0.128513   phase=  -57.416 deg.
-----
calsinu:
spin de Y+DY       amplitude=  0.126794   phase=   31.534 deg.
-----
calsinu:
spin de Z          amplitude=  0.002689   phase=   61.343 deg.
-----
calsinu:
spin de X+DX-Sin   amplitude=  0.005941   phase=   31.472 deg.
-----
calsinu:
spin de Y+DY-Sin   amplitude=  0.003865   phase=  146.796 deg.
-----
deconvo:
deconvolution des 3 formes d'onde- t.f. directe
      fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre      0.100 et   11.500 Hz
correction du retard d'echantillonnage de la voie 1
      fft inverse effectuee
      fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre      0.100 et   11.500 Hz
correction du retard d'echantillonnage de la voie 2
      fft inverse effectuee
      fft directe effectuee
correction des gains d'antenne effectuee en module et en phase
filtrage de la forme d'onde entre      0.100 et   11.500 Hz
correction du retard d'echantillonnage de la voie 3
      fft inverse effectuee
t.f. inverse- retour au domaine temps
-----
calsinu:
spin de -X+DX decon. amplitude=  2.476671   phase=  -65.693 deg.
-----
calsinu:
spin de -Y+DY decon. amplitude=  2.791475   phase=   29.321 deg.
-----
calsinu:
spin de -Z+0 decon. amplitude=  1.696915   phase=  -22.704 deg.
-----
etab3:

13) ULF+Di-Fs nT, SC spinning system
resolution (s):  0.0430003
Nb. de courbes:   3
Nb. de points : 1024
```

rotfix:

changement de base sur les formes d'onde, repere final=repere sat. fixe
x meridien rotdeg= 143.4 nbp= 1024 fs= 0.16981

etab3:

14) ULF+Despins nT, SC fixed system with X in geo. meridian (Z14=Z13)
resolution (s): 0.0430003
Nb. de courbes: 3
Nb. de points : 1024

etab3:

15) continuous mag. field, nT, SC fixed system with X in geo. meridian
resolution (s): 5.5040398
Nb. de courbes: 3
Nb. de points : 8

etab3:

16) continuous mag. field, nT, VDH system
resolution (s): 5.5040398
Nb. de courbes: 3
Nb. de points : 8

=====
efinpro: fin normale de caliulf //////////////////////////////////////
=====

Annexe 3: Example of caliulf.resu file

GEOS S300 ULF Data - Results of Caliulf.exe

Input Parameters:

28 08 1978 09 14 40 40 1024 0.1 11.5 0 *

GEOS:

2

Number of sequences:

1

Status of selected panels :

1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16

1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1

Number of panels:

16

Starting date (njul50,iday,imonth,iyear)

10466 28 8 1978

Starting Time (ih, im, is, ims)

9 14 40 862

Antenna direction at t=0 in VDH system:

V	D	H	r	theta	phi
-0.80313	-0.59580	-0.00226	1.000	90.130	-143.430
0.59580	-0.80312	-0.00393	1.000	90.225	-53.430
0.00052	-0.00451	0.99999	1.000	0.259	-83.370

Spin frequency (Hz) and Spin period (s) :

0.169815 5.88877

Axes of geo. meridian system in VDH:

V	D	H	r	theta	phi
1.00000	0.00000	-0.00052	1.000	90.030	0.000
0.00000	0.99999	0.00451	1.000	89.742	90.000
0.00052	-0.00451	0.99999	1.000	0.259	-83.370

1) XYZ Telemetry data

resolution (s):

0.0430003

Number of curves:

3

Number of points :

1024

151.000000 85.000000 104.000000

148.000000 98.000000 98.000000

149.000000 106.000000 105.000000

154.000000 108.000000 126.000000

..... total of 3*1024 data

2) Gains XY and Z

resolution (s):

1.3760099

Number of curves:

2

Number of points :

32

25.000000 25.000000

25.000000 25.000000

..... total of 2*32 data

3) XYZ volt, SC spinning system

resolution (s):

0.0430003

Number of curves:

3

Number of points :

1024

-0.020800 0.032000 0.016800

-0.018400 0.021600 0.021600

..... total of 3*1024 data

4) XY volt, SC fixed system with X in geo. meridian
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
0.035771 -0.013307
0.027911 -0.005110
..... total of 2*1024 data

5) DX,DY volt, SC spinning system
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
-0.098000 0.040000
-0.094000 0.046000
..... total of 2*1024 data

6) DX,DY volt, SC fixed system with X in geo. meridian
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
0.102539 0.026264
0.101919 0.023761
..... total of 2*1024 data

7) X+DX,Y+DY volt, SC spinning system
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
-0.118800 0.072000
-0.112400 0.067600
..... total of 2*1024 data

8) X+DX,Y+DY volt, SC fixed system with X in geo. meridian
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
0.138310 0.012956
0.129829 0.018651
..... total of 2*1024 data

9) XYZ nT, SC spinning system
resolution (s):
0.0430003
Number of curves:
3
Number of points :
1024
-9.363985 0.214724 3.385161
-9.253702 0.451304 3.339678
..... total of 3*1024 data

10) XYZ nT, SC fixed system with X in geo. meridian
resolution (s):
0.0430003
Number of curves:
2
Number of points :
1024
7.648420 5.406662
7.456457 5.498721
..... total of 2*1024 data

11) XYZ nT, VDH system
resolution (s):
0.0430003
Number of curves:
3
Number of points :
1024
7.650204 5.391349 3.405490
7.458218 5.483612 3.360522
..... total of 3*1024 data

12) XYZ nT, XYB strength line system
resolution (s):
0.0430003
Number of curves:
3
Number of points :
1024
7.606159 -6.375210 -0.832280
7.422702 -6.431157 -0.729290
..... total of 3*1024 data

13) ULF+Di-Fs nT, SC spinning system
resolution (s):
0.0430003
Number of curves:
3
Number of points :
1024
-8.553311 5.523035 3.581485
-8.330996 5.702818 3.535944
..... total of 3*1024 data

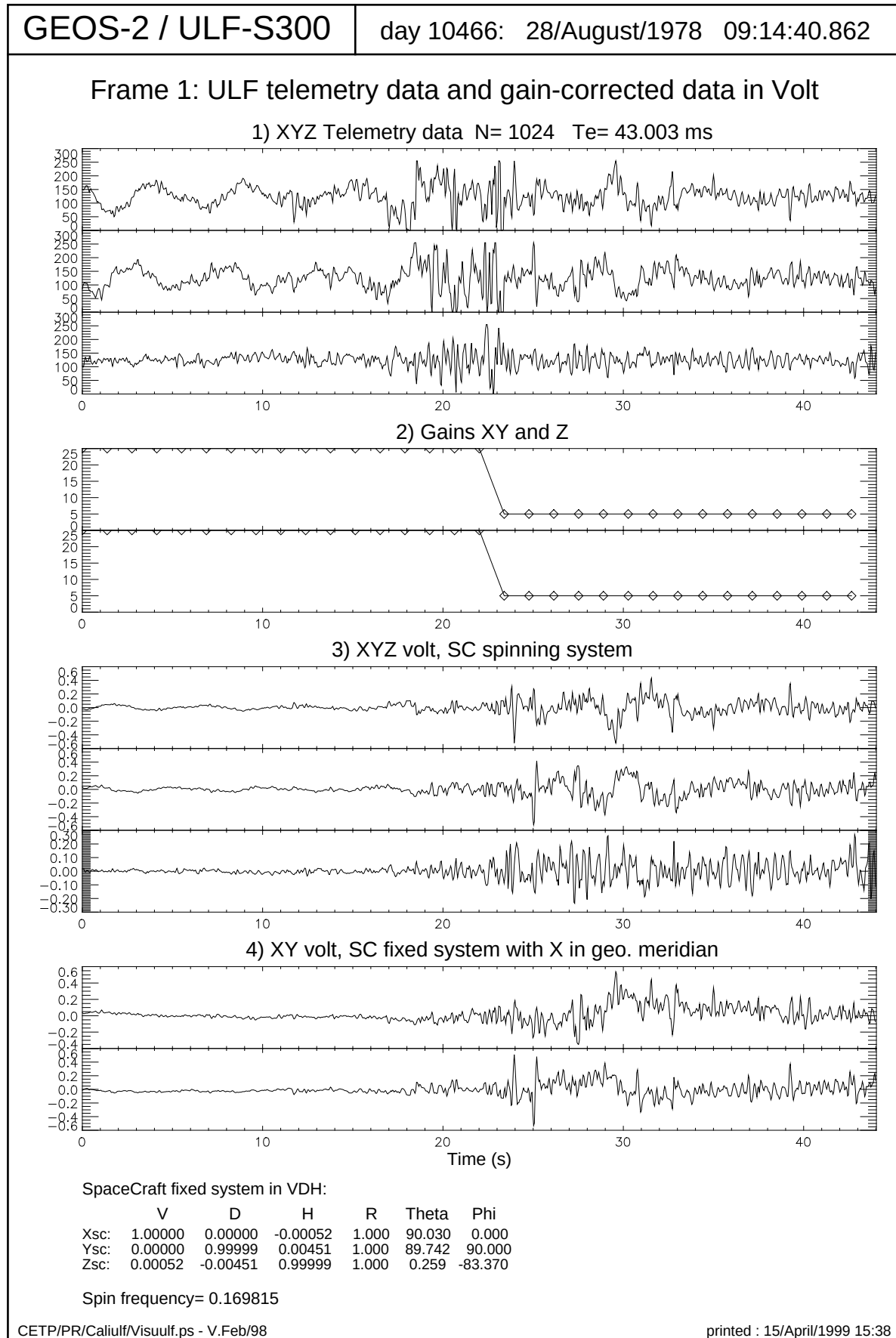
14) ULF+Despins nT, SC fixed system with X in geo. meridian (Z14=Z13)
resolution (s):
0.0430003
Number of curves:
3
Number of points :
1024
7.063542 56.933361 3.581485
6.963902 57.118820 3.535944
..... total of 3*1024 data

15) continuous mag. field, nT, SC fixed system with X in geo. meridian
resolution (s):
5.5040398
Number of curves:
3
Number of points :
8
-10.069319 57.180893 -96.203194
-3.720074 47.024006 -97.774605
..... total of 3*8 data

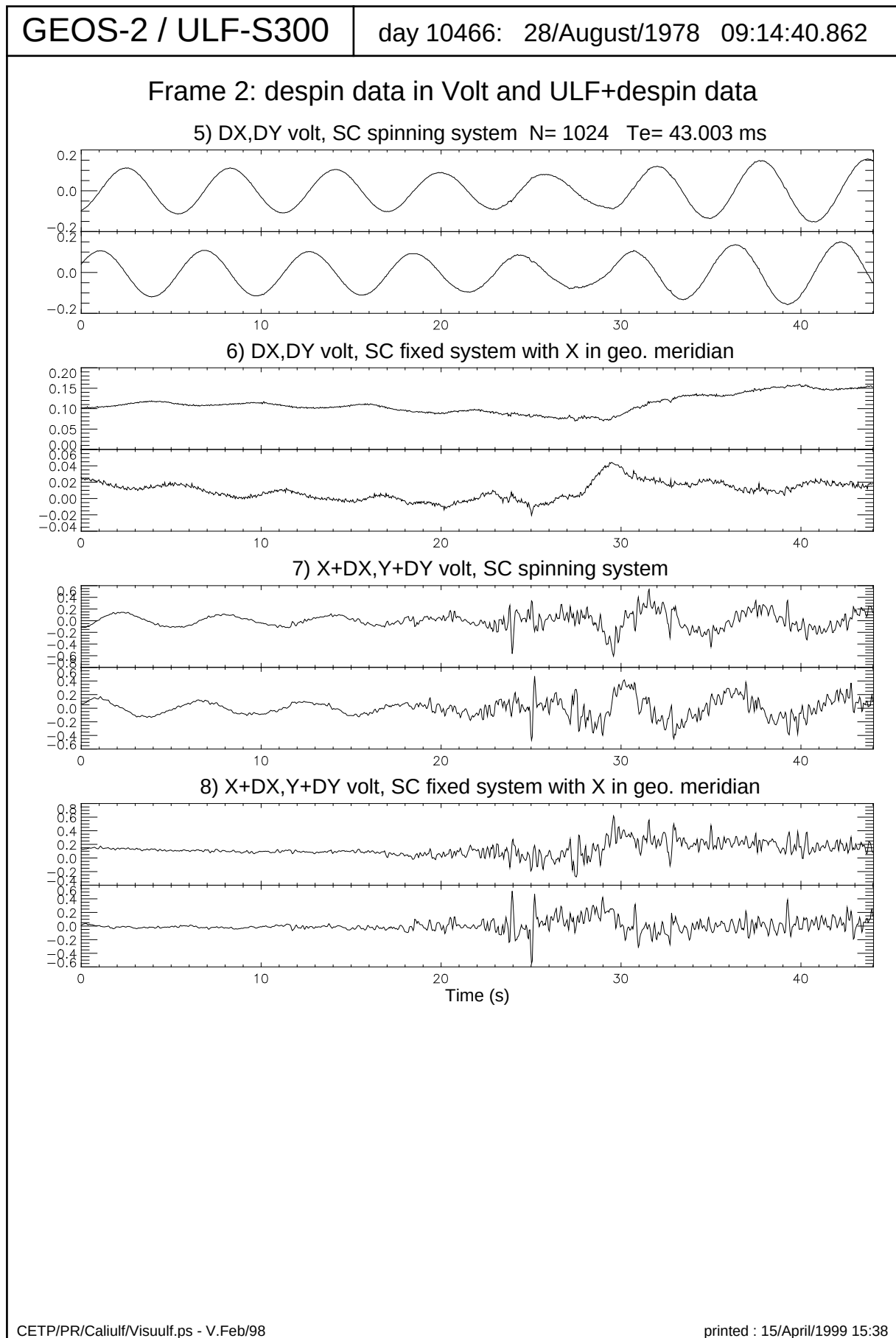
16) continuous mag. field, nT, VDH system
resolution (s):
5.5040398
Number of curves:
3
Number of points :
8
-10.000000 57.700001 -95.900002
-3.600000 47.599998 -97.500000
..... total of 3*8 data

Annexe 4: Example of ULF visualisation data

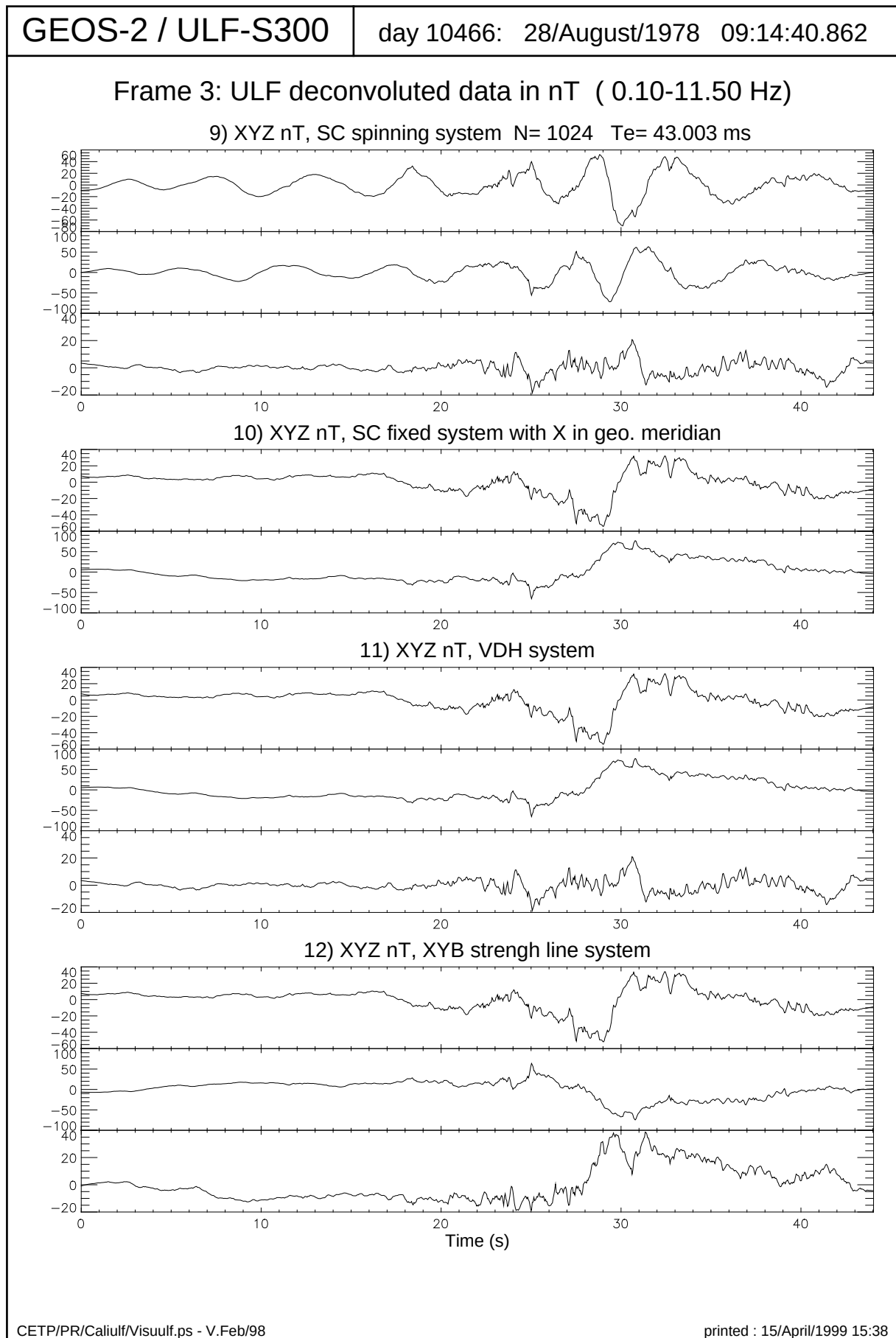
A4-a) Example of PostScript file corresponding to frame 1 (visuulf_s1_f1.ps):



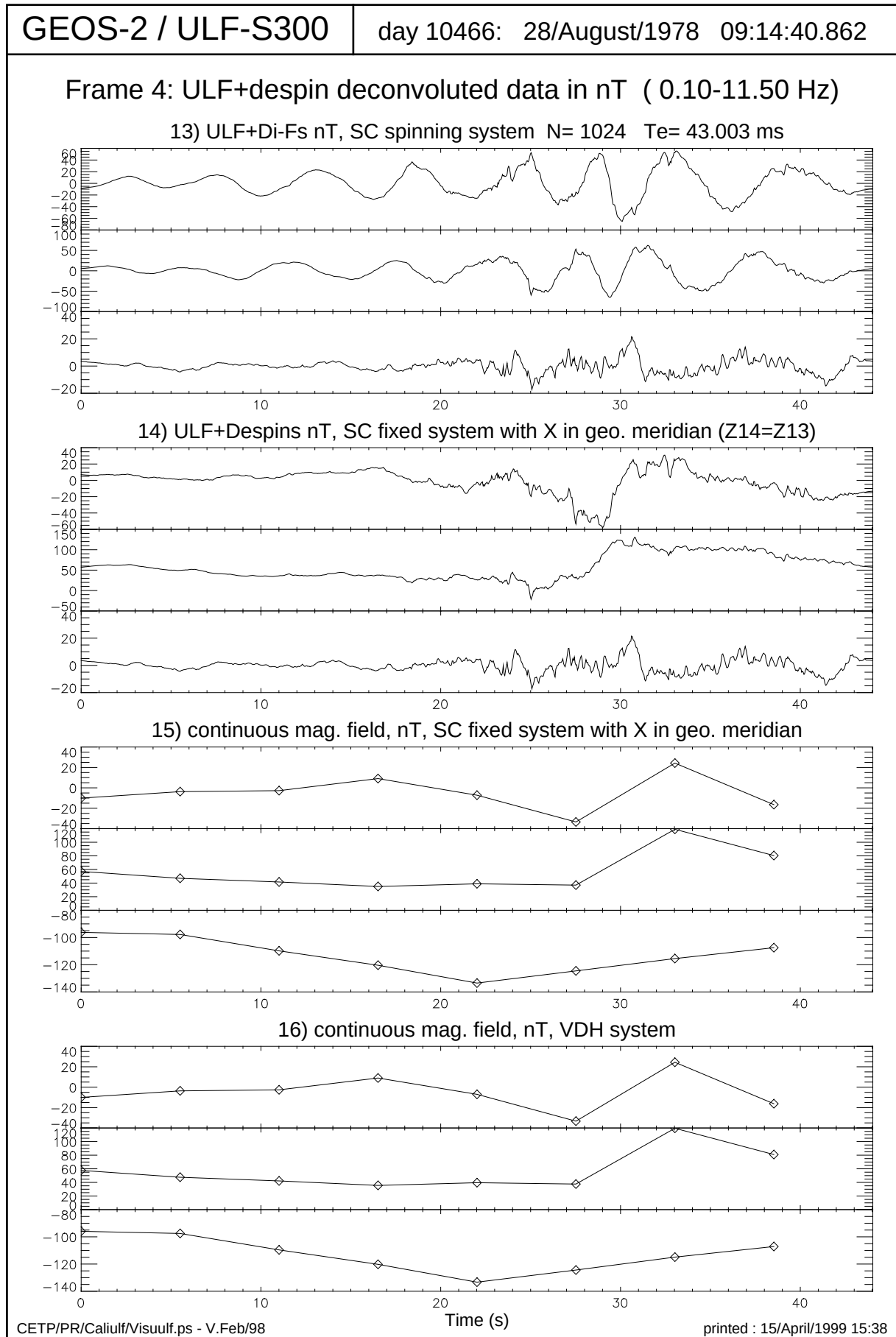
A4-b) Example of PostScript file corresponding to frame 2 (visuulf_s1_f2.ps):



A4-c) Example of PostScript file corresponding to frame 3 (visuulf_s1_f3.ps):



A4-d) Example of PostScript file corresponding to frame 4 (visuulf_s1_f4.ps):



Annexe 5: List of ULF files available from CDPF

The name of the GEOS-1 and GEOS-2 ULF data files available from the CDPF server are given below. The name contains the date and time of the beginning data period, for instance the UBF_19770427_164141 files correspond to the data period from 1977/04/27-16:41:41. The complete period is explained beside each name, and the size of each file, in Kbytes, is also given. The total size of the data sets are about 2.312 Go for GEOS-1 and 17.097 Go for GEOS-2.

A5-a) GEOS-1 Data Files

File names of a data set
Data set name : UBF_GEOS1
Ultra Low Frequency magnetic field data from GEOS1

Name	From	To	Size (Kb)
UBF_19770427_164141	1977/04/27-16:41:41	1977/05/06-16:15:55	26464
UBF_19770506_161030	1977/05/06-16:10:30	1977/05/12-19:02:24	29037
UBF_19770512_185659	1977/05/12-18:56:59	1977/05/18-21:08:30	29218
UBF_19770519_125416	1977/05/19-12:54:16	1977/05/24-18:43:04	28122
UBF_19770524_183739	1977/05/24-18:37:39	1977/05/31-15:06:18	28945
UBF_19770531_150053	1977/05/31-15:00:53	1977/06/06-14:56:38	30833
UBF_19770606_145113	1977/06/06-14:51:13	1977/06/12-14:46:57	30309
UBF_19770612_144133	1977/06/12-14:41:33	1977/06/17-19:34:34	30275
UBF_19770618_105338	1977/06/18-10:53:38	1977/06/22-18:25:42	25910
UBF_19770622_182017	1977/06/22-18:20:17	1977/06/28-18:16:04	29260
UBF_19770628_181039	1977/06/28-18:10:39	1977/07/03-18:37:21	27063
UBF_19770703_183157	1977/07/03-18:31:57	1977/07/08-18:58:39	29416
UBF_19770708_185314	1977/07/08-18:53:14	1977/07/13-16:23:50	29729
UBF_19770713_161825	1977/07/13-16:18:25	1977/07/18-16:45:08	29503
UBF_19770718_163943	1977/07/18-16:39:43	1977/07/23-17:06:26	30924
UBF_19770723_170101	1977/07/23-17:01:01	1977/07/28-17:27:45	30896
UBF_19770728_172220	1977/07/28-17:22:20	1977/08/02-11:56:46	26112
UBF_19770802_115122	1977/08/02-11:51:22	1977/08/07-18:10:20	29747
UBF_19770807_180455	1977/08/07-18:04:55	1977/08/13-12:08:24	30085
UBF_19770813_120259	1977/08/13-12:02:59	1977/08/14-16:36:02	8569
UBF_19770814_142810	1977/08/14-14:28:10	1977/08/19-09:02:38	29024
UBF_19770819_085713	1977/08/19-08:57:13	1977/08/23-09:54:53	26152
UBF_19770823_094929	1977/08/23-09:49:29	1977/08/27-16:39:25	30004
UBF_19770827_163400	1977/08/27-16:34:00	1977/09/01-14:04:34	30897
UBF_19770901_135910	1977/09/01-13:59:10	1977/09/06-11:29:45	30408
UBF_19770906_112420	1977/09/06-11:24:20	1977/09/12-14:16:13	25620
UBF_19770912_141048	1977/09/12-14:10:48	1977/09/17-14:37:31	31063
UBF_19770917_143206	1977/09/17-14:32:06	1977/09/22-14:58:49	29410
UBF_19770922_145325	1977/09/22-14:53:25	1977/09/27-06:31:44	27303
UBF_19770927_062619	1977/09/27-06:26:19	1977/10/02-12:45:17	29959
UBF_19771002_123952	1977/10/02-12:39:52	1977/10/07-04:18:13	25513
UBF_19771007_041248	1977/10/07-04:12:48	1977/10/11-12:32:38	30485
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A5-b) GEOS-2 Data Files

File names of a data set

Data set name : UBF_GEOS2

Ultra Low Frequency magnetic field data from GEOS2

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UBF_19790721_234932	1979/07/21-23:49:32	1979/07/24-01:49:10	31783

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UBF_19790728_062042	1979/07/28-06:20:42	1979/07/30-08:20:20	31750
UBF_19790730_081456	1979/07/30-08:14:56	1979/08/01-10:14:34	31664
UBF_19790801_100909	1979/08/01-10:09:09	1979/08/03-12:08:47	31695
UBF_19790803_120322	1979/08/03-12:03:22	1979/08/05-14:03:01	31656
UBF_19790805_135736	1979/08/05-13:57:36	1979/08/07-15:57:14	31857
UBF_19790807_155149	1979/08/07-15:51:49	1979/08/09-17:51:27	31836
UBF_19790809_174603	1979/08/09-17:46:03	1979/08/11-19:45:41	31781
UBF_19790811_194016	1979/08/11-19:40:16	1979/08/13-21:39:54	31725
UBF_19790813_213430	1979/08/13-21:34:30	1979/08/15-23:34:08	31666
UBF_19790815_232843	1979/08/15-23:28:43	1979/08/18-01:28:21	31113
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UBF_19790820_031710	1979/08/20-03:17:10	1979/08/22-05:16:48	31700
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UBF_19790824_070537	1979/08/24-07:05:37	1979/08/26-09:05:15	31765
UBF_19790826_085950	1979/08/26-08:59:50	1979/08/28-10:59:29	31529
UBF_19790828_105404	1979/08/28-10:54:04	1979/08/30-12:53:42	31723
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UBF_19790901_151430	1979/09/01-15:14:30	1979/09/03-17:14:08	31789
UBF_19790903_170844	1979/09/03-17:08:44	1979/09/05-20:47:26	31208
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UBF_19790914_041854	1979/09/14-04:18:54	1979/09/16-06:18:32	31682
UBF_19790916_061307	1979/09/16-06:13:07	1979/09/18-08:12:45	31698
UBF_19790918_080720	1979/09/18-08:07:20	1979/09/20-10:06:58	31438
UBF_19790920_100133	1979/09/20-10:01:33	1979/09/22-12:01:11	31673
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UBF_19791024_151058	1979/10/24-15:10:58	1979/10/26-17:10:37	31809
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UBF_19810619_204940	1981/06/19-20:49:40	1981/06/23-19:03:34	31857
UBF_19810623_185809	1981/06/23-18:58:09	1981/06/27-18:36:03	31857
UBF_19810627_183038	1981/06/27-18:30:38	1981/07/01-19:25:36	31857
UBF_19810701_192011	1981/07/01-19:20:11	1981/07/05-07:06:51	31857

UBF_19810705_070126	1981/07/05-07:01:26	1981/07/08-20:25:54	31857
UBF_19810708_202029	1981/07/08-20:20:29	1981/07/12-18:24:04	31857
UBF_19810712_181839	1981/07/12-18:18:39	1981/07/15-18:12:52	31857
UBF_19810715_180728	1981/07/15-18:07:28	1981/07/19-16:08:50	31857
UBF_19810719_160325	1981/07/19-16:03:25	1981/07/23-16:41:29	31857
UBF_19810723_163604	1981/07/23-16:36:04	1981/07/26-16:41:19	31857
UBF_19810726_163554	1981/07/26-16:35:54	1981/07/30-03:29:05	31857
UBF_19810730_032340	1981/07/30-03:23:40	1981/08/02-13:36:04	31857
UBF_19810802_133039	1981/08/02-13:30:39	1981/08/06-00:20:38	31857
UBF_19810806_001514	1981/08/06-00:15:14	1981/08/09-22:07:55	31857
UBF_19810809_220230	1981/08/09-22:02:30	1981/08/12-22:18:40	31857
UBF_19810812_221315	1981/08/12-22:13:15	1981/08/16-19:38:05	31857
UBF_19810816_193240	1981/08/16-19:32:40	1981/08/20-17:22:55	31857
UBF_19810820_171730	1981/08/20-17:17:30	1981/08/23-19:37:30	31850
UBF_19810823_193205	1981/08/23-19:32:05	1981/08/27-17:48:59	31857
UBF_19810827_174334	1981/08/27-17:43:34	1981/08/31-22:59:57	31857
UBF_19810831_225432	1981/08/31-22:54:32	1981/09/05-00:40:12	31857
UBF_19810905_003447	1981/09/05-00:34:47	1981/09/08-02:45:52	31719
UBF_19810908_024028	1981/09/08-02:40:28	1981/09/13-00:42:01	31857
UBF_19810913_003636	1981/09/13-00:36:36	1981/09/16-15:41:09	31857
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UBF_19810920_193433	1981/09/20-19:34:33	1981/09/25-16:45:37	31857
UBF_19810925_164012	1981/09/25-16:40:12	1981/09/28-12:19:53	31857
UBF_19810928_121429	1981/09/28-12:14:29	1981/10/02-19:25:49	31857
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UBF_19811006_033151	1981/10/06-03:31:51	1981/10/10-20:24:04	31857
UBF_19811010_201839	1981/10/10-20:18:39	1981/10/14-07:12:32	31857
UBF_19811014_070707	1981/10/14-07:07:07	1981/10/17-18:10:39	31857
UBF_19811017_180514	1981/10/17-18:05:14	1981/10/21-17:34:28	31857
UBF_19811021_172904	1981/10/21-17:29:04	1981/10/24-17:52:05	31857
UBF_19811024_174640	1981/10/24-17:46:40	1981/10/29-16:33:20	31857
UBF_19811029_162755	1981/10/29-16:27:55	1981/11/02-04:13:27	31857
UBF_19811102_040802	1981/11/02-04:08:02	1981/11/06-02:16:29	31857
UBF_19811106_021104	1981/11/06-02:11:04	1981/11/09-02:16:41	31857
UBF_19811109_021116	1981/11/09-02:11:16	1981/11/12-22:19:51	31857
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UBF_19811116_213145	1981/11/16-21:31:45	1981/11/19-22:15:05	31857
UBF_19811119_220941	1981/11/19-22:09:41	1981/11/23-21:14:36	31857
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UBF_19811127_194935	1981/11/27-19:49:35	1981/11/30-19:59:26	31857
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UBF_19811204_184650	1981/12/04-18:46:50	1981/12/07-18:58:26	31857
UBF_19811207_185301	1981/12/07-18:53:01	1981/12/11-18:45:12	31857
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UBF_19811215_170105	1981/12/15-17:01:05	1981/12/18-17:21:14	31857
UBF_19811218_171549	1981/12/18-17:15:49	1981/12/23-05:03:32	31857
UBF_19811223_045807	1981/12/23-04:58:07	1981/12/27-03:14:11	31857
UBF_19811227_030847	1981/12/27-03:08:47	1981/12/30-03:28:21	31857
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UBF_19820110_083056	1982/01/10-08:30:56	1982/01/13-19:53:48	31857
UBF_19820113_194823	1982/01/13-19:48:23	1982/01/17-17:53:03	31857
UBF_19820117_174738	1982/01/17-17:47:38	1982/01/20-19:22:07	31857
UBF_19820120_191643	1982/01/20-19:16:43	1982/01/24-18:12:27	31857
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UBF_19820212_043027	1982/02/12-04:30:27	1982/02/13-05:10:04	8493
UBF_19820216_030904	1982/02/16-03:09:04	1982/02/19-04:22:18	31857
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UBF_19820423_033427	1982/04/23-03:34:27	1982/04/26-04:39:35	31857
UBF_19820426_043411	1982/04/26-04:34:11	1982/04/30-03:03:03	31857
UBF_19820430_025739	1982/04/30-02:57:39	1982/05/04-01:48:14	31857
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UBF_19820514_221820	1982/05/14-22:18:20	1982/05/18-20:13:56	31857
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UBF_19820525_195440	1982/05/25-19:54:40	1982/05/29-19:35:43	31857
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UBF_19820608_214531	1982/06/08-21:45:31	1982/06/12-23:12:30	31857
UBF_19820612_230706	1982/06/12-23:07:06	1982/06/16-11:39:10	31857
UBF_19820616_113345	1982/06/16-11:33:45	1982/06/20-00:06:59	31857
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UBF_19820630_162801	1982/06/30-16:28:01	1982/07/02-19:47:18	31857
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UBF_19820722_185711	1982/07/22-18:57:11	1982/07/25-00:50:15	31857
UBF_19820725_004450	1982/07/25-00:44:50	1982/07/27-04:47:39	31857
UBF_19820729_145719	1982/07/29-14:57:19	1982/07/31-17:10:09	31857
UBF_19820731_170445	1982/07/31-17:04:45	1982/08/02-21:03:00	31857
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UBF_19820827_082512	1982/08/27-08:25:12	1982/08/29-12:11:37	31857
UBF_19820829_120612	1982/08/29-12:06:12	1982/09/01-21:17:10	31857
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UBF_19820905_121736	1982/09/05-12:17:36	1982/09/09-17:08:47	31857
UBF_19820909_170322	1982/09/09-17:03:22	1982/09/12-15:17:52	31857
UBF_19820912_151228	1982/09/12-15:12:28	1982/09/16-07:22:28	31857
UBF_19820916_071703	1982/09/16-07:17:03	1982/09/19-16:51:27	31855
UBF_19820919_164603	1982/09/19-16:46:03	1982/09/23-13:17:12	31578
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UBF_19820926_040314	1982/09/26-04:03:14	1982/09/29-13:24:48	31857
UBF_19821003_014330	1982/10/03-01:43:30	1982/10/06-12:48:41	31857
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UBF_19821010_003446	1982/10/10-00:34:46	1982/10/12-02:44:17	31857
UBF_19821014_063002	1982/10/14-06:30:02	1982/10/16-09:21:56	31857
UBF_19821016_091631	1982/10/16-09:16:31	1982/10/18-20:54:18	31857
UBF_19821018_204853	1982/10/18-20:48:53	1982/10/21-01:06:59	31857
UBF_19821021_010134	1982/10/21-01:01:34	1982/10/23-07:46:13	31857
UBF_19821023_074048	1982/10/23-07:40:48	1982/10/25-09:43:00	31857
UBF_19821027_124327	1982/10/27-12:43:27	1982/10/29-14:59:47	31857
UBF_19821029_145422	1982/10/29-14:54:22	1982/10/31-16:56:17	31857
UBF_19821031_165052	1982/10/31-16:50:52	1982/11/02-23:33:56	31857
UBF_19821107_070757	1982/11/07-07:07:57	1982/11/09-09:12:42	31857
UBF_19821109_090718	1982/11/09-09:07:18	1982/11/11-12:28:31	31857
UBF_19821115_183117	1982/11/15-18:31:17	1982/11/17-20:48:34	31857

UBF_19821117_204309	1982/11/17-20:43:09	1982/11/19-22:55:28	31857
UBF_19821119_225003	1982/11/19-22:50:03	1982/11/22-00:49:54	31857
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UBF_19821124_023921	1982/11/24-02:39:21	1982/11/26-04:41:51	31857
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UBF_19821215_165959	1982/12/15-16:59:59	1982/12/17-19:05:41	31857
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UBF_19830110_234117	1983/01/10-23:41:17	1983/01/13-02:47:47	31857
UBF_19830117_100033	1983/01/17-10:00:33	1983/01/19-13:57:51	32332
UBF_19830119_135226	1983/01/19-13:52:26	1983/01/21-17:12:36	31857
UBF_19830121_170711	1983/01/21-17:07:11	1983/01/23-20:02:04	31857
UBF_19830123_195639	1983/01/23-19:56:39	1983/01/25-22:17:07	31857
UBF_19830125_221143	1983/01/25-22:11:43	1983/01/28-00:14:15	31857
UBF_19830128_000850	1983/01/28-00:08:50	1983/01/30-02:31:40	31857
UBF_19830130_022615	1983/01/30-02:26:15	1983/01/31-22:02:46	26584

Annexe 6: Example of caliaux.out file

GEOS-1 / GEOS-2

CALIAUX V 2.0

Decommutation et calibration des donnees auxiliaires

P.Robert, CETP, Fevrier 1996
Version pour le CDDP, Aout 1999

lfauxin:

Nom du fichier de donnees auxiliaires ?
/data1/ida/robert/geos/geosdata/DAVL-GEOS1

Date (jj,mm,aa) ? [27,04,1977]
21 12 1977

Heure de debut (hh,mn,ss) ? [15,00,00]
0 1 30

Duree de la periode (jours) ? [3]
3.00000

resume des parametres d'entree:
21 12 1977 00 01 30 3.00

soit 2943 blocs de 88.0646 sec.

calgeos:

GEOS-1/2: GEOS-1

sensrot:

sens de rotation= proton

posfic: positionnement sur fichier de type davl

debut fichier : njul50,ih,im,is = 9978 14 56 56

fichier correspondant a GEOS-1
nombre de records: 399240
taille en octets : 25551360

fin fichier : = 10400 23 10 20
demande : = 10216 0 1 30
fait : njul50,ih,im,is,imil= 10216 0 1 21 774

eentaux:

Donnees Auxiliaires GEOS - Resultats de caliaux.exe
parametre d'entree: 21 12 1977 00 01 30 3.00
GEOS: 1
Nombre de blocs : 2943

ecdavl: premier bloc

nrec	nJul	millisec	num lsf	lat	long	R km	sunsat	rasc	decli	Qua
1	10216	81774	22346944	20.65	54.40	32750.	51.24	0.00	3.93	0
	SpinX	SpinY	SpinZ	Dip_X	Dip_Y	Dip_Z	Sun_X	Sun_Y	Sun_Z	
	-0.0013	-0.0998	-0.9948	0.2409	-0.1628	0.9567	-0.6259	0.7564	-0.1891	
3	10216	258305	22347072	constime: milt,idif=			257905		-400	
4	10216	345974	22347136	constime: milt,idif=			346369		395	

83	10216	7303078	22352192	constime: milt, idif=	7303173	95
84	10216	7391240	22352256	constime: milt, idif=	7391142	-98
359	10216	31609089	22369856	constime: milt, idif=	31609164	75
390	10216	34339168	22371840	constime: milt, idif=	34339073	-95
936	10216	82422994	22406784	constime: milt, idif=	82422112	-882
940	10216	82774368	22407040	constime: milt, idif=	82775258	890
941	10216	82863326	22407104	constime: milt, idif=	82862432	-894
944	10216	83126624	22407296	constime: milt, idif=	83127523	899
950	10216	83655923	22407680	constime: milt, idif=	83655008	-915
1337	10217	31337134	22432448	constime: milt, idif=	31337228	94
1350	10217	32482068	22433280	constime: milt, idif=	32481966	-102
1925	10217	83118405	22470080	constime: milt, idif=	83118868	463
2299	10218	29654769	22494016	constime: milt, idif=	29654870	101
2310	10218	30623581	22494720	constime: milt, idif=	30623473	-108
2901	10218	82670318	22532544	constime: milt, idif=	82669405	-913

ecdav1: dernier bloc

nrec	nJul	millisec	num lsf	lat	long	R km	sunsat	rasc	decli	Qua
2943	10218	86369096	22535232	21.43	54.09	33600.	50.06	0.00	3.78	0

SpinX	SpinY	SpinZ	Dip_X	Dip_Y	Dip_Z	Sun_X	Sun_Y	Sun_Z
-0.0197	-0.1146	-0.9930	0.2547	-0.1634	0.9530	-0.6419	0.7463	-0.1751

=====
efinpro: fin normale de caliaux` //////////////////////////////////////
=====

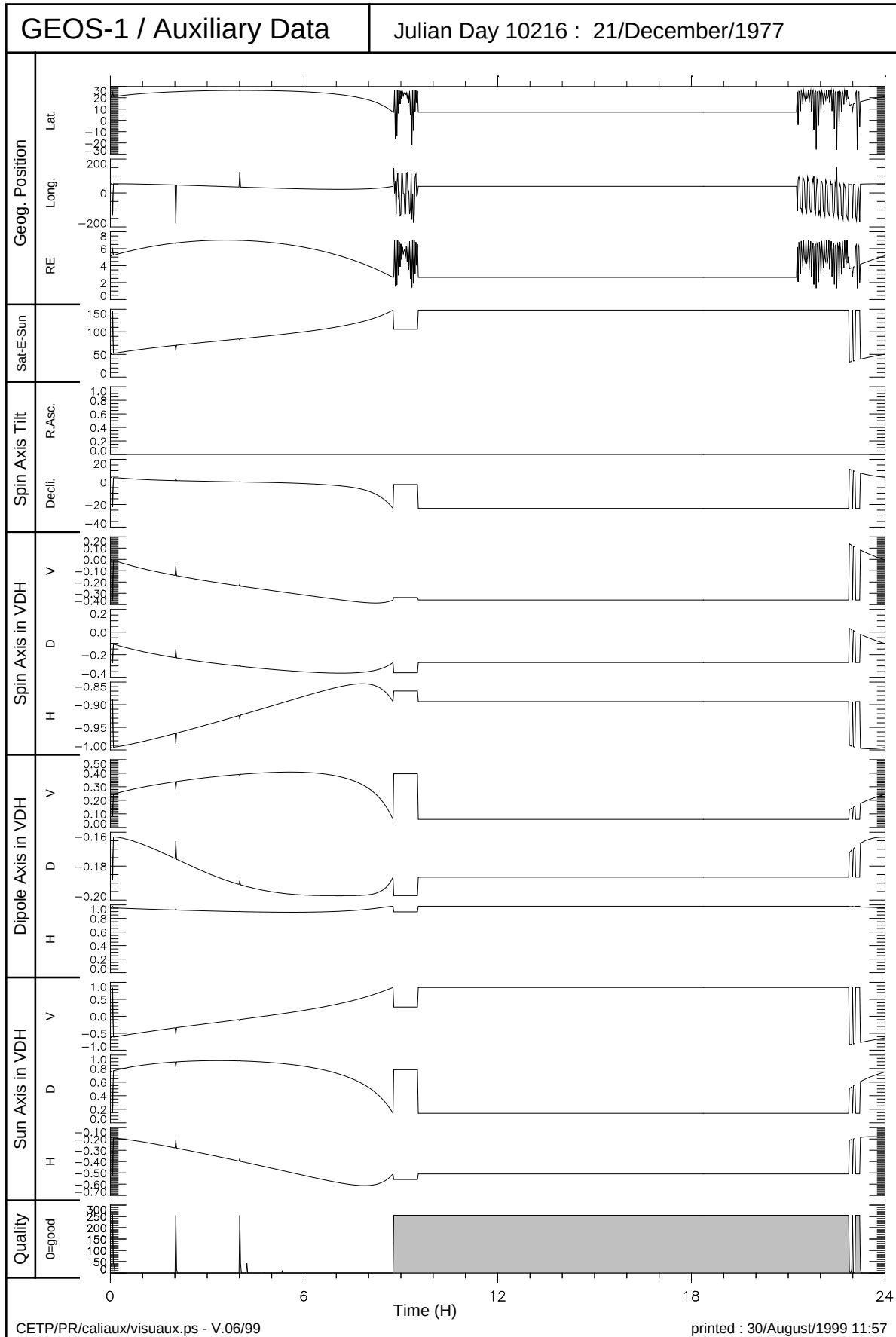
Annexe 7: Example of caliaux.resu file

```
GEOS Auxiliary Data - Results of caliaux.exe
Input Parameters:
21 12 1977 00 01 30      3.00
GEOS:
  1
Number of blocs:
  2943
Starting date (njul50, iday, imonth, iyear)
10216 21 12 1977
Starting Time (ih, im, is, ims)
  0 1 21 774
nrec njul millisec num lsf lat long R km sunsat rasc decli Qua
  SpinX SpinY SpinZ Dip_X Dip_Y Dip_Z Sun_X Sun_Y Sun_Z
1 10216 81774 22346944 20.65 54.40 32750. 51.24 0.00 3.93 0
-0.0013 -0.0998 -0.9948 0.2409 -0.1628 0.9567 -0.6259 0.7564 -0.1891
2 10216 169841 22347008 20.75 54.40 32930. 51.55 0.00 3.90 0
-0.0036 -0.1020 -0.9946 0.2425 -0.1628 0.9563 -0.6217 0.7597 -0.1898
3 10216 258305 22347072 24.76 -129.57 37700. 146.93 0.00 -22.16 255
-0.3713 -0.2742 -0.8868 0.0842 -0.1880 0.9785 0.8379 0.1430 -0.5265
4 10216 345974 22347136 20.99 54.38 33380. 52.34 0.00 3.82 38
-0.0095 -0.1077 -0.9940 0.2467 -0.1628 0.9552 -0.6108 0.7680 -0.1916
5 10216 434040 22347200 21.07 54.37 33520. 52.58 0.00 3.79 27
-0.0113 -0.1095 -0.9938 0.2480 -0.1628 0.9549 -0.6074 0.7706 -0.1922
6 10216 522105 22347264 21.12 54.36 33620. 52.75 0.00 3.77 0
-0.0126 -0.1107 -0.9936 0.2489 -0.1628 0.9547 -0.6051 0.7723 -0.1926
7 10216 610171 22347328 21.22 54.34 33790. 53.05 0.00 3.73 0
-0.0148 -0.1128 -0.9933 0.2505 -0.1629 0.9542 -0.6010 0.7753 -0.1933
8 10216 698237 22347392 21.31 54.32 33960. 53.33 0.00 3.69 0
-0.0170 -0.1149 -0.9931 0.2521 -0.1629 0.9538 -0.5970 0.7782 -0.1940
9 10216 786303 22347456 21.40 54.29 34120. 53.62 0.00 3.64 0
-0.0192 -0.1169 -0.9928 0.2536 -0.1630 0.9534 -0.5930 0.7811 -0.1948
10 10216 874369 22347520 21.49 54.26 34290. 53.90 0.00 3.60 0
-0.0214 -0.1190 -0.9925 0.2552 -0.1630 0.9530 -0.5891 0.7839 -0.1955

. . . . .

2931 10218 85312303 22534464 20.18 54.09 31450. 46.34 0.00 4.72 0
0.0090 -0.0878 -0.9959 0.2339 -0.1634 0.9583 -0.6902 0.7032 -0.1698
2932 10218 85400369 22534528 20.30 54.11 31630. 46.67 0.00 4.63 0
0.0064 -0.0902 -0.9957 0.2358 -0.1633 0.9579 -0.6861 0.7072 -0.1701
2933 10218 85488435 22534592 20.41 54.13 31820. 46.99 0.00 4.54 0
0.0039 -0.0925 -0.9955 0.2376 -0.1633 0.9574 -0.6820 0.7110 -0.1704
2934 10218 85576501 22534656 20.52 54.14 32010. 47.31 0.00 4.46 0
0.0014 -0.0948 -0.9953 0.2394 -0.1633 0.9570 -0.6779 0.7149 -0.1707
2935 10218 85664567 22534720 20.63 54.15 32190. 47.63 0.00 4.37 0
-0.0009 -0.0971 -0.9951 0.2412 -0.1632 0.9566 -0.6738 0.7186 -0.1711
2936 10218 85752633 22534784 20.74 54.15 32370. 47.94 0.00 4.29 0
-0.0033 -0.0994 -0.9949 0.2430 -0.1632 0.9561 -0.6697 0.7223 -0.1715
2937 10218 85840699 22534848 20.84 54.16 32550. 48.25 0.00 4.21 0
-0.0057 -0.1017 -0.9946 0.2447 -0.1632 0.9557 -0.6657 0.7259 -0.1719
2938 10218 85928766 22534912 20.94 54.15 32730. 48.56 0.00 4.14 0
-0.0081 -0.1039 -0.9944 0.2464 -0.1632 0.9552 -0.6617 0.7295 -0.1724
2939 10218 86016832 22534976 21.04 54.15 32910. 48.86 0.00 4.06 0
-0.0105 -0.1061 -0.9941 0.2481 -0.1633 0.9548 -0.6577 0.7330 -0.1729
2940 10218 86104898 22535040 21.14 54.14 33080. 49.17 0.00 3.99 0
-0.0128 -0.1082 -0.9939 0.2498 -0.1633 0.9543 -0.6537 0.7364 -0.1734
2941 10218 86192964 22535104 21.24 54.13 33260. 49.47 0.00 3.92 0
-0.0151 -0.1104 -0.9936 0.2514 -0.1633 0.9539 -0.6497 0.7398 -0.1740
2942 10218 86281030 22535168 21.33 54.11 33430. 49.76 0.00 3.85 0
-0.0174 -0.1125 -0.9933 0.2531 -0.1633 0.9535 -0.6458 0.7431 -0.1745
2943 10218 86369096 22535232 21.43 54.09 33600. 50.06 0.00 3.78 0
-0.0197 -0.1146 -0.9930 0.2547 -0.1634 0.9530 -0.6419 0.7463 -0.1751
```

Annexe 8: Example of Auxiliary visualisation data



NOTES
